Part two: Kai Schneider

A review on wavelet transforms and their applications to MHD and plasma turbulence II

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In collaboration with Marie Farge, ENS Paris

Methods for Analyzing Turbulence Data Meudon, 29 May 2015

Fast visible light camera

A fast camera from the Nancy team (G. Bonhomme and F. Brochard) was installed on Tore-Supra (N. Fedorczak and P. Monier-Garbet).

An helical Abel transform relates the plasma light emissivity S to the integral of the volume emissivity received by the camera I=KS, where K is a compact continuous operator. Reconstruction of S from I is an inverse problem which becomes very difficult when S is corrupted by noise, then solving K⁻¹ is an ill-posed problem.

> Tomographic inversion using wavelet-vaguelette decomposition as an alternative to SVD (Singular Value Decomposition).

Image tomography



Tomography inversion in presence of noise

Image received by the camera: integral of the volume emissivity I=KS



Nguyen, Fedorczak, Brochard, Bonhomne, Schneider, Farge, Monier-Garbet, Nuclear Fusion, **52**, 2012

Denoised plasma emissivity



Movie from a fast camera in Tore-Supra tokamak



Noisy images

Coherent structures

Incoherent background

Nguyen, Fedorczak, Brochard, Bonhomne, Schneider, Farge, Monier-Garbet, Nuclear Fusion, **52**, 2012

Noise reduction in plasma simulations using particles

- Accuracy of particle simulations is limited by noise (statistical sampling, not enough particles and grid effects)
- Wavelet based density estimation, accurate estimation of distribution functions with localized sharp features
- Preservation of moments in the distribution functions
- No a priori selection of a global smoothing scale
- No constraints on the dimensionality
- · Computationally efficient: same order as for finite size particle approach

Nguyen van yen, del-Castillo-Negrete, Schneider, Farge and Chen, J. Comput. Phys., **229**, 2010

Noise reduction in plasma simulations using particles



Noise reduction in plasma simulations using particles

Collisional guiding center transport data (Delta5d)



RMS error estimate with respect to the reference density computed with $N_p = 1024 \times 10^3$. Error reduction by about a factor 2.

Particle in wavelets scheme for Vlasov-Poisson equation

- Plasma distribution function is discretized using tracer particles
- The charge distribution is reconstructed using wavelet based density estimation
- Wavelet expansion of the Dirac delta functions corresponding to each particle
- Wavelet Galerkin Poisson solver to compute the electric potential from the electron charge density (diagonal preconditioning)
- Improvement of precision compared to a classical PIC scheme for a given number of particles

Nguyen van yen, Sonnendrücker, Schneider and Farge, ESAIM Proc., **32**, 2011

Particle in wavelets scheme for Vlasov-Poisson equation

Two-stream instability test case



Particle distribution function at t=10 (left) and t=30 (right).

Nguyen van yen, Sonnendrücker, Schneider and Farge, ESAIM Proc., **32**, 2011 L^2 error on the electric field at t = 30, as a function of number of particles



Papers on applications to tokamaks http://wavelets.ens.fr

Marie Farge, Kai Schneider and Pascal Devynck, 2006 Extraction of coherent bursts in turbulent edge plasma using orthogonal wavelets *Physics of Plasmas*, **13**(2), 042304, 1-11

Romain Nguyen van yen, Diego del Castilo–Negrete, Kai Schneider, Marie Farge and Guangye Chen, 2010 Wavelet–based density estimation for noise reduction in plasma simulation using particles *J. Comput. Phys.*, **229**(8), 2821-2839

Romain Nguyen van yen, Eric Sonnendrücker, Kai Schneider and Marie Farge, 2011 Particle-in-wavelet scheme for the 1D Vlasov-Poisson equations ESAIM Proc., **32**, 134-148

Romain Nguyen van yen, Nicolas Fedorczak, Frédéric Brochard, Kai Schneider, Marie Farge and Pascale Monier-Garbet, 2012 Tomographic reconstruction of tokamak edge turbulence light emission from single image using wavelet-vaguelette decomposition *Nuclear Fusion, IAEA (International Atomic Energy Agency),* **52**, 013005, 1-11

Coherent structures extraction in 3D MHD flow



Velocity



Magnetic field





Current density



Coherent Vorticity Simulation (CVS)



Coherent Vortex Simulation (CVS)



Schneider & Farge, 2000, Comp. Rend. Acad. Sci. Paris, 328

- 1. Selection of the wavelet coefficients whose modulus is larger than the threshold.
- 2. Construction of a 'graded-tree' which defines the 'interface' between the coherent and incoherent wavelet coefficients.
- 3. Addition of a 'security zone' which corresponds to dealiasing.

Schneider & Farge, 2002, Appl. Comput. Harmonic Anal., 12

Schneider, Farge et al., 2005, J. Fluid Mech., 534(5)

3D turbulent mixing layer



3D turbulent mixing layer



3D turbulent mixing layer



Adaptive computation using wavelets

Koster, Schneider, Griebel, Farge Numerical Flow Simulation II, **75**, Springer, 2001

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Adaptive computation using wavelets



Roussel and Schneider, **75**, 2000



Intermittency and geometrical statistics of 3d homogeneous magnetohydrodynamic turbulence using wavelets

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> Turbulence workshop 2015: Analyzing turbulence data May 29, 2015, Observatoire de Paris, Meudon, France



Motivation

Different wavelet based tools to examine scale-dependent statistics.

What are similarities and differences of small-scale intermittency in HD and MHD turbulence?

Here we use and generalize the diagnostics introduced in Yoshimatsu et al., Phys. Rev. E, 79, 2009, which are based on the orthonormal wavelet decomposition.



Orthogonal wavelets



Ref.: M. Farge. Annu. Rev. Fluid Mech., 24, 1992 K. Schneider, O. Vasilyev. Annu. Rev. Fluid Mech., 42, 2010



Wavelet decomposition

The vector field v, having a mean value $\langle v \rangle$ (which vanishes in the present applications for all components), can be decomposed into an orthogonal wavelet series

$$\boldsymbol{v}(\boldsymbol{x}) = \langle \boldsymbol{v} \rangle + \sum_{j=0}^{J-1} \boldsymbol{v}_j(\boldsymbol{x}),$$
 (1)

where v_j is the contribution of v at scale 2^{-j} defined by

$$\boldsymbol{v}_{j}(\boldsymbol{x}) = \sum_{\mu=1}^{7} \sum_{i_{1},i_{2},i_{3}=0}^{2^{j}-1} \tilde{\boldsymbol{v}}_{\mu,\lambda} \psi_{\mu,\lambda}(\boldsymbol{x}).$$
(2)

with the wavelet coefficients $ilde{m{v}}_{\mu,\lambda}=\langlem{v},\psi_{\mu,\lambda}
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Scale dependent statistics (I)

Scale dependent energy

$$e_j^v = \langle oldsymbol{v}_j, oldsymbol{v}_j
angle /2$$
 with $ar{E^v} = \sum_{j=0}^{J-1} e_j^v$

Wavelet energy spectrum

$$\tilde{E}_{j}^{v} = \frac{1}{\Delta k_{j}} \langle e_{j}^{v,\ell} \rangle_{c} \quad \text{with } k_{j} = k_{\psi} 2^{j} \text{ and } k_{\psi} = 0.77$$
$$\Delta k_{j} = k_{j} \ln 2$$

Spatial variability of the energy spectrum

$$\tilde{\sigma}_{j}^{v} = \frac{1}{\Delta k_{j}} \sqrt{\langle (e_{j}^{v,\ell})^{2} \rangle_{c} - (\langle e_{j}^{v,\ell} \rangle_{c})^{2}}$$



Scale dependent statistics (II)

Scale dependent flatness

$$F[\boldsymbol{v}_j] = \frac{\langle (v_j^\ell)^4 \rangle_c}{\{\langle (v_j^\ell)^2 \rangle_c\}^2}$$

Link with wavelet energy spectrum and its standard deviation

$$F[\boldsymbol{v}_j] = \left(\frac{\tilde{\sigma}_j^v}{\tilde{E}_j^v}\right)^2 + 1$$

Ref.: W. Bos, L. Liechtenstein and K. Schneider. Phys. Rev. E, 76, 2007.



Scale dependent statistics (III)

Scale dependent helicites (kinetic, cross, magnetic)

$$H_j^{\mathrm{K}}(\boldsymbol{x}) = \boldsymbol{u}_j \cdot \boldsymbol{\omega}_j \quad H_j^{\mathrm{C}}(\boldsymbol{x}) = \boldsymbol{u}_j \cdot \boldsymbol{b}_j \quad H_j^{\mathrm{M}}(\boldsymbol{x}) = \boldsymbol{a}_j \cdot \boldsymbol{b}_j$$

Scale dependent relative helicites

$$h_j^{\mathrm{K}}(\boldsymbol{x}) = \frac{H_j^{\mathrm{K}}}{|\boldsymbol{u}_j||\boldsymbol{\omega}_j|} \qquad h_j^{\mathrm{C}}(\boldsymbol{x}) = \frac{H_j^{\mathrm{C}}}{|\boldsymbol{u}_j||\boldsymbol{b}_j|} \qquad h_j^{\mathrm{M}}(\boldsymbol{x}) = \frac{H_j^{\mathrm{M}}}{|\boldsymbol{a}_j||\boldsymbol{b}_j|}$$



DNS of MHD Turbulence

Forced 3D incompressible MHD turbulence without mean magnetic field in a 2π periodic box Ω .

$$\partial_t \boldsymbol{u} + (\boldsymbol{u} \cdot \nabla) \boldsymbol{u} = -\frac{1}{\rho_0} \nabla P + \boldsymbol{j} \times \boldsymbol{b} + \nu \Delta \boldsymbol{u} + \boldsymbol{f},$$

 $\partial_t \boldsymbol{b} + (\boldsymbol{u} \cdot \nabla) \boldsymbol{b} = (\boldsymbol{b} \cdot \nabla) \boldsymbol{u} + \eta \Delta \boldsymbol{b},$
 $\nabla \cdot \boldsymbol{u} = 0,$
 $\nabla \cdot \boldsymbol{b} = 0,$

where *t* is time, *f* is an external force, *P* is the pressure, ν is the kinematic viscosity, η is the magnetic diffusivity, and $\partial_t = \partial/\partial t$. The Prandtl number Pr is set to 1, i.e., $\eta = \nu$.



Vorticity and current density

\bar{E}^{u}	\bar{E}^b	$\bar{H}^{\rm K}(\times 10^{-3})$	$\bar{H}^{C}(\times 10^{-3})$	$ar{H}^{\mathrm{M}}$	$\eta_{\rm IK}(\times 10^{-3})$	R^u_λ	R^b_λ
0.238	0.618	7.16	-5.77	0.503	8.79	150	306



Resolution N= 512^3

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Wavelet mean kinetic and magnetic energy spectra





Spatial variability of the kinetic and magnetic energy spectra





Scale dependent PDFs







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Scale dependent PDFs of relative kinetic helicity

MHD turbulence



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Scale dependent PDFs of relative kinetic helicity

HD turbulence



Scale dependent PDFs of relative cross helicity

MHD turbulence



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Scale dependent PDFs of relative magnetic helicity

MHD turbulence





Eulerian and Lagrangian acceleration

$$oldsymbol{a}^E = -(oldsymbol{u} \cdot
abla)oldsymbol{u} - rac{1}{
ho_0}
abla P + oldsymbol{j} imes oldsymbol{b} +
u
abla^2 oldsymbol{u},$$

 $oldsymbol{a}^L = -rac{1}{
ho_0}
abla P + oldsymbol{j} imes oldsymbol{b} +
u
abla^2 oldsymbol{u},$



Scale dependent PDFs of Eulerian acceleration





Scale dependent PDFs of Lagrangian acceleration





Scale dependent flatness of Eulerian and Lagrangian acceleration





Eulerian and Lagrangian time derivatives of the magnetic field

 $\partial_t \boldsymbol{b}$ and $D_t \boldsymbol{b}$

 $D_t \boldsymbol{b} = \partial_t \boldsymbol{b} + (\boldsymbol{u} \cdot \nabla) \boldsymbol{b}$



Scale dependent PDFs of Eulerian time derivative





Scale dependent PDFs of Lagrangian time derivative



092304-8 Yoshimatsu et al.



Scale dependent flatness of Eulerian and Lagrangian time derivatives





Conclusions (I)

- Geometrical and scale-dependent statistics 3D MHD turbulence using orthogonal wavelets
- Wavelet decomposition yields clear scale separation spatial localization to quantify the intermittency of the flow
- Appplication to DNS data of stat. stat. MHD turbulence at R_{λ}^{u} = 150

Results:

- Magnetic field more intermittent than velocity (faster increase of flatness with scale)
- Multiscale measures to study geometrical statistics:
 - Relative scale-dependent kinetic, cross and magnetic helicities.
 - Higher probability for velocity and vorticity vectors to be aligned or anti-aligned, i.e., helical flow, at small scales for MHD turbulence.



Conclusions (II)

Analysis of scale-dependent statistics of Eulerian and Lagrangian accelerations, and corresponding time-derivatives of the magnetic field.

Different dynamics of MHD compared to HD turbulent flows. In MHD turbulence, intermittency of the Lagrangian acc. is comparable to that of the Eulerian acc..

In HD turbulence, the Lagrangian acc. exhibits substantially stronger intermittency than the Eulerian one.

The Eulerian time-derivative of the magnetic field is more intermittent than the corresponding Lagrangian time derivative.

Intermittency in MHD turbulence is different from HD turbulence.

Ref.:K. Yoshimatsu, K. Schneider, N. Okamoto, Y. Kawahura and M. Farge. Intermittency and geometrical statistics of three-dimensional homogeneous magnetohydrodynamic turbulence: A wavelet viewpoint. *Phys. Plasmas*, 18, 092304, 2011.

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