

# 2-D hybrid simulations of kinetic plasma turbulence

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and

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*Franci et al., ApJ Letters 2015 (arXiv:1503.05457)*

*Franci et al., to be submitted*

# Several numerical approaches:

- Hall MHD: Matthaeus 2003,...
- Electron MHD: Meyrand&Galtier 2013
- Gyrokinetics: Howes et al 2008, 2011
- Landau Fluid: Passot et al 2014
- Hybrid-PIC: Markoski&Vasquez 2011
- Hybrid-Vlasov: Servidio 2012, Perrone 2013, Valentini 2014
- Full PIC: Camporeale&Burgess 2011
- ...

# Hybrid approximation

Kinetic ions (PIC) and fluid electrons

- MHD ✓
- ion scales ✓
- electron scales ✗

isothermal electrons:

$$P_e = nk_B T_e$$

physical limit:

$$kd_e = 1 \text{ or } kd_p \sim 40$$

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Hall-MHD

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# Simulation Setup

Units and normalizations:

space:  $d_p = c/\omega_p = V_A/\Omega_p$

time:  $\Omega_p^{-1}$

magnetic field:  $B_0$

velocity:  $V_A$

2048x2048 grid

$\Delta t = 0.025$ ,  $\Delta x = \Delta y = 0.125$

( $0.02 < k < 20$  in Fourier space)

$\beta_p = \beta_e = 0.5$

$T_{\perp}/T_{\parallel} = 1$

**$B_0$**  out of simulation plane

Run	$\Delta x/d_p$	$L_{\text{box}}/d_p$	$\eta$	ppc
A	0.125	256	$5 \times 10^{-4}$	8000
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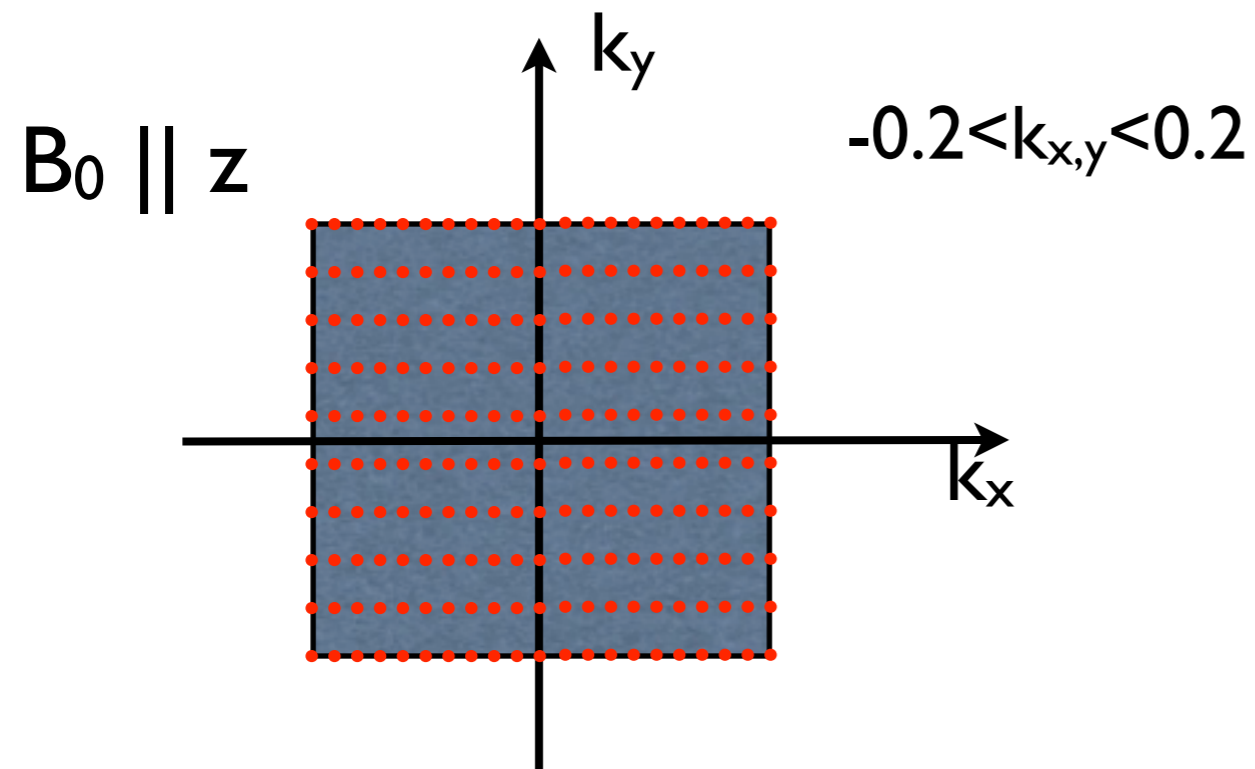
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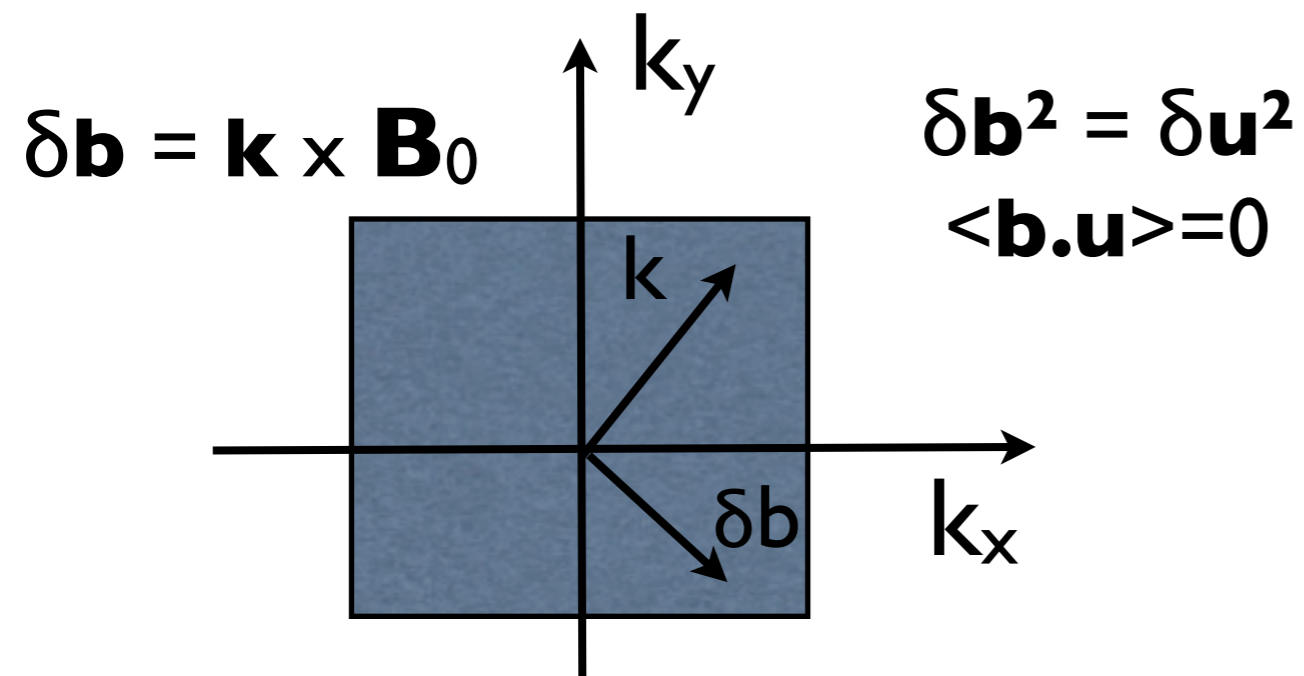
$\sim 3 \times 10^{10}$  in total!

# Initialization of in plane fluctuations

k-vector orthogonal to the main field

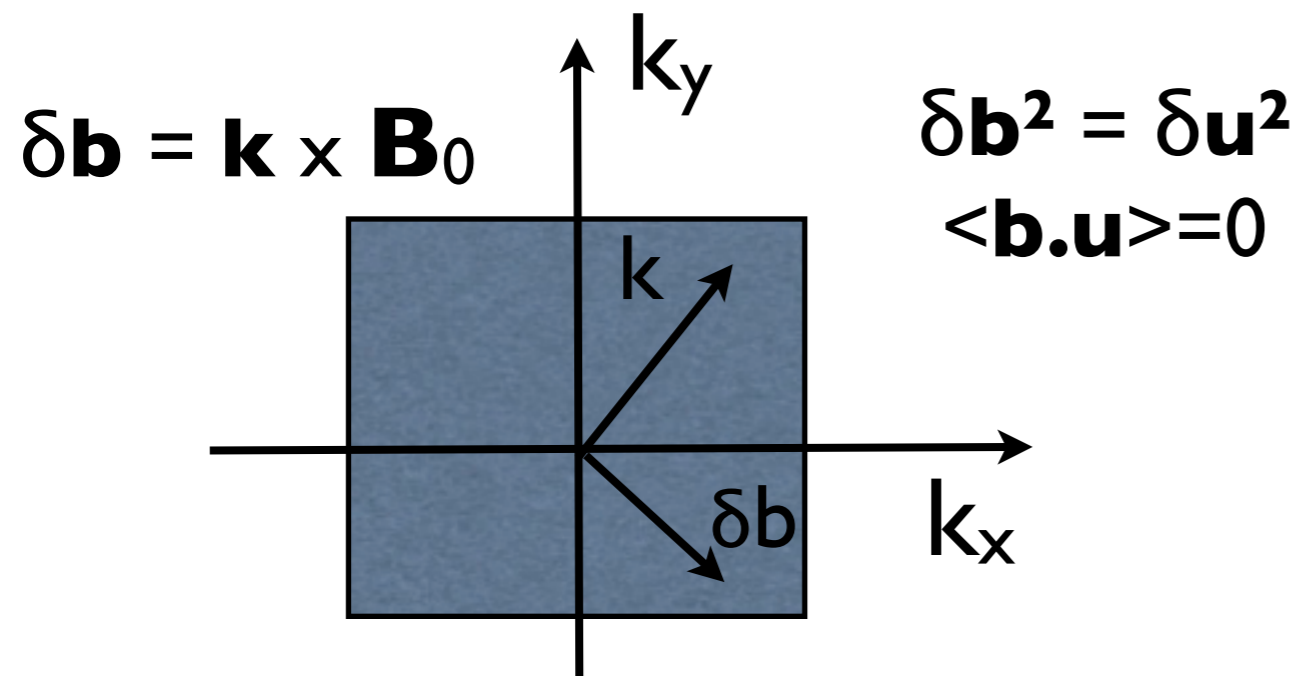
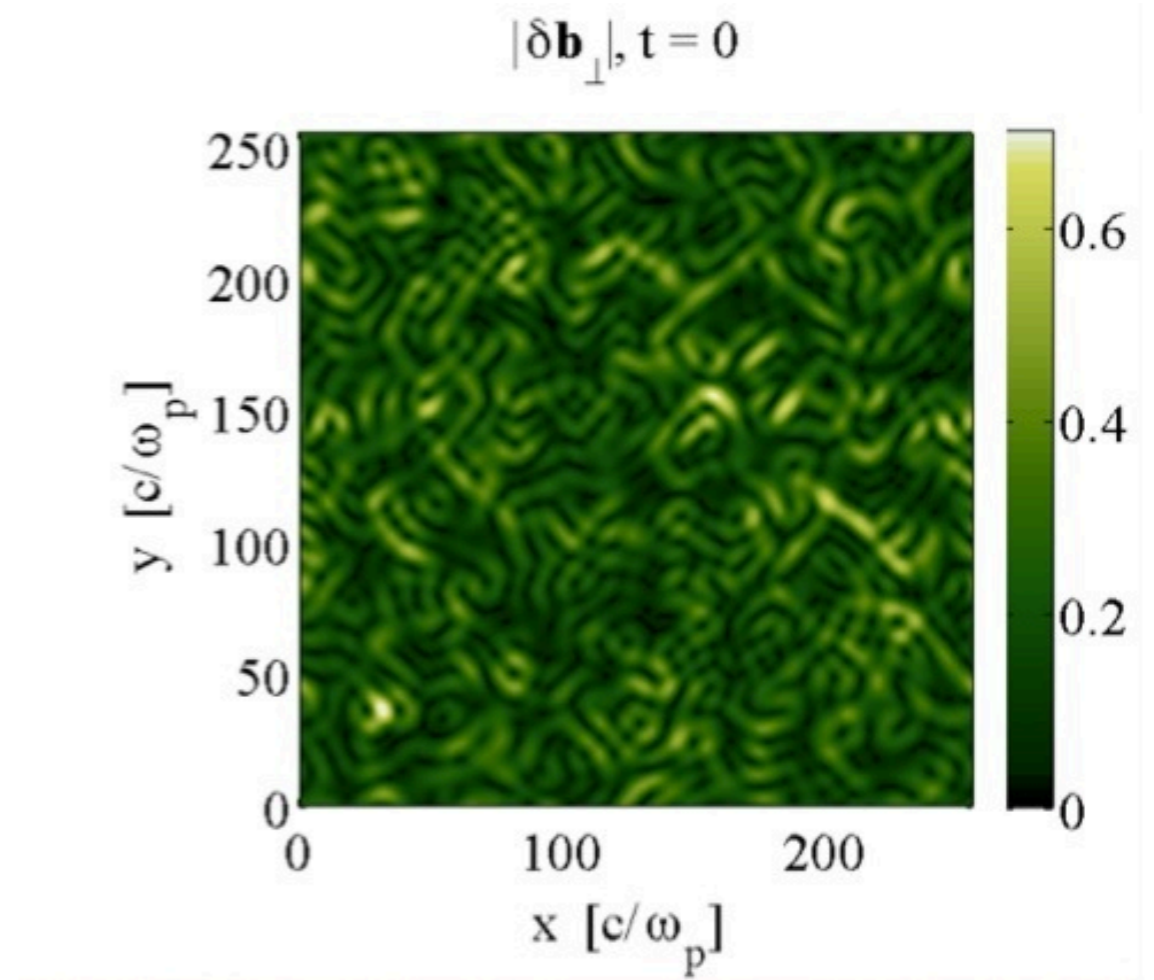
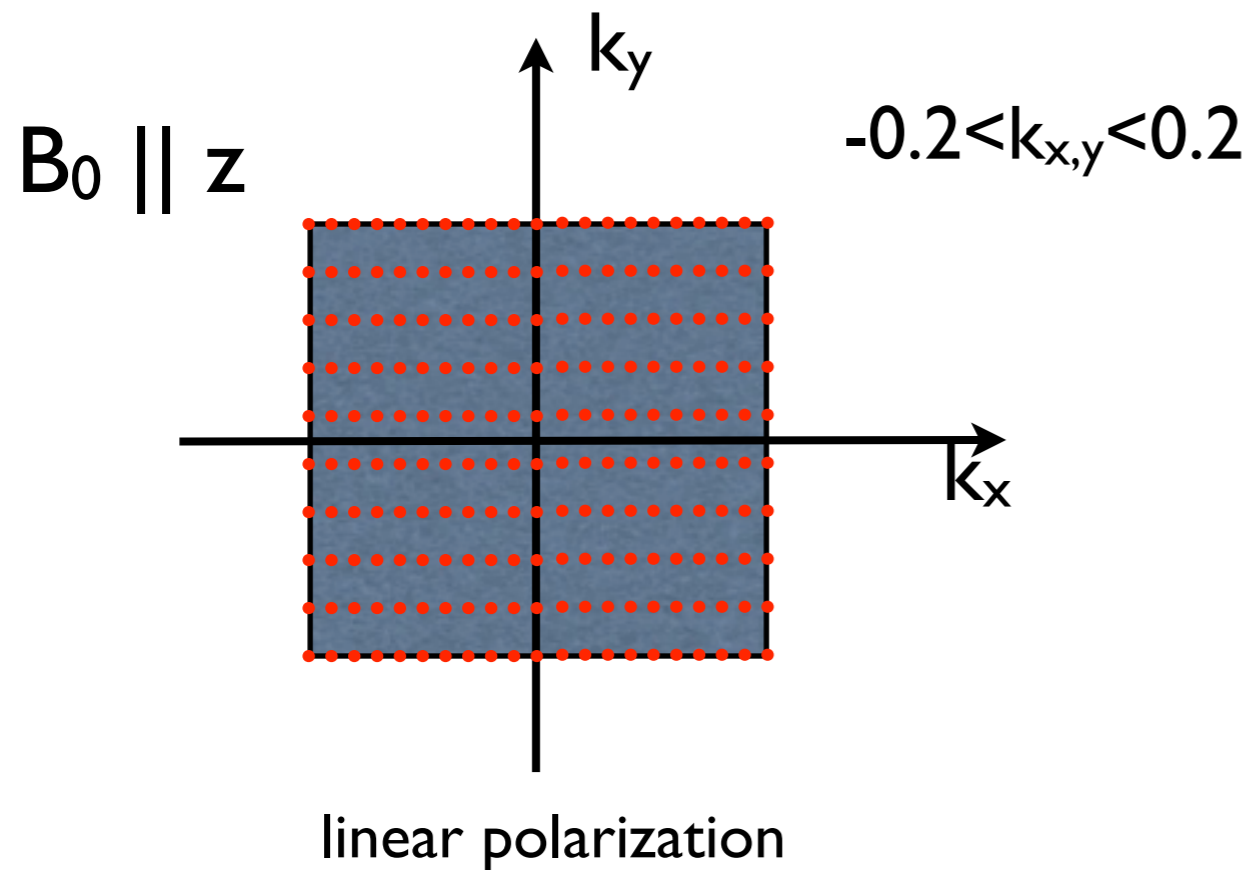


linear polarization



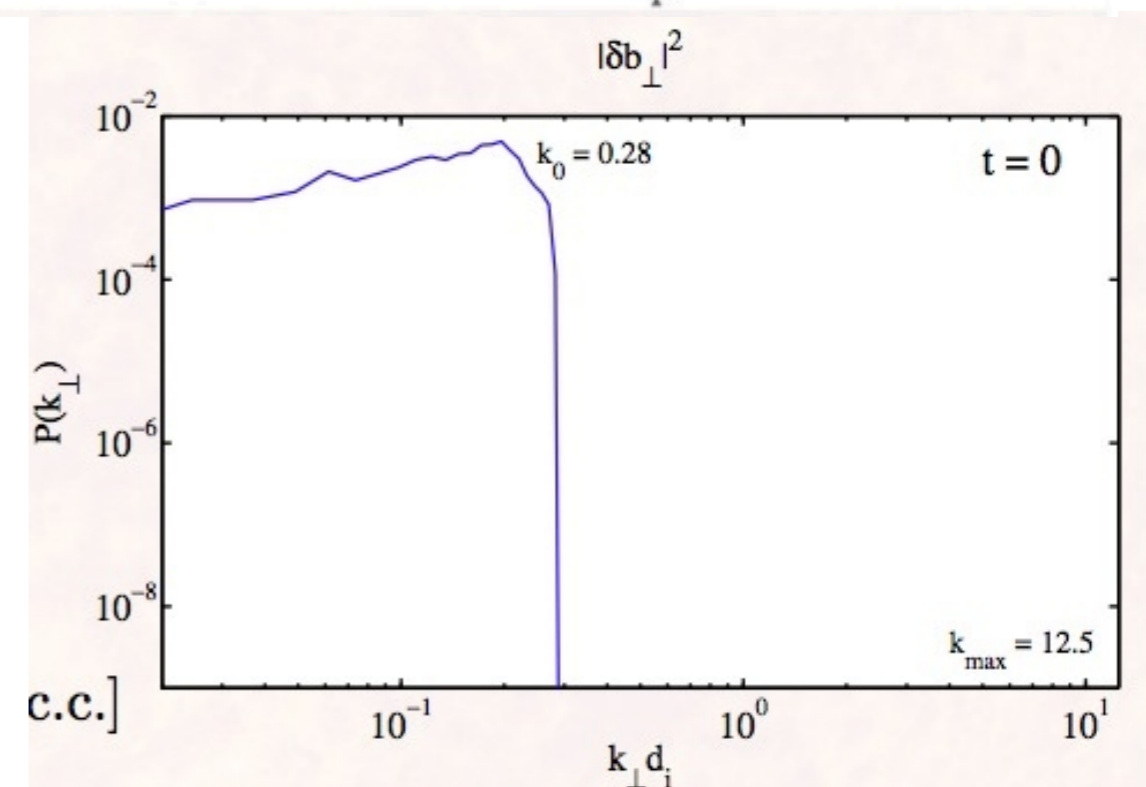
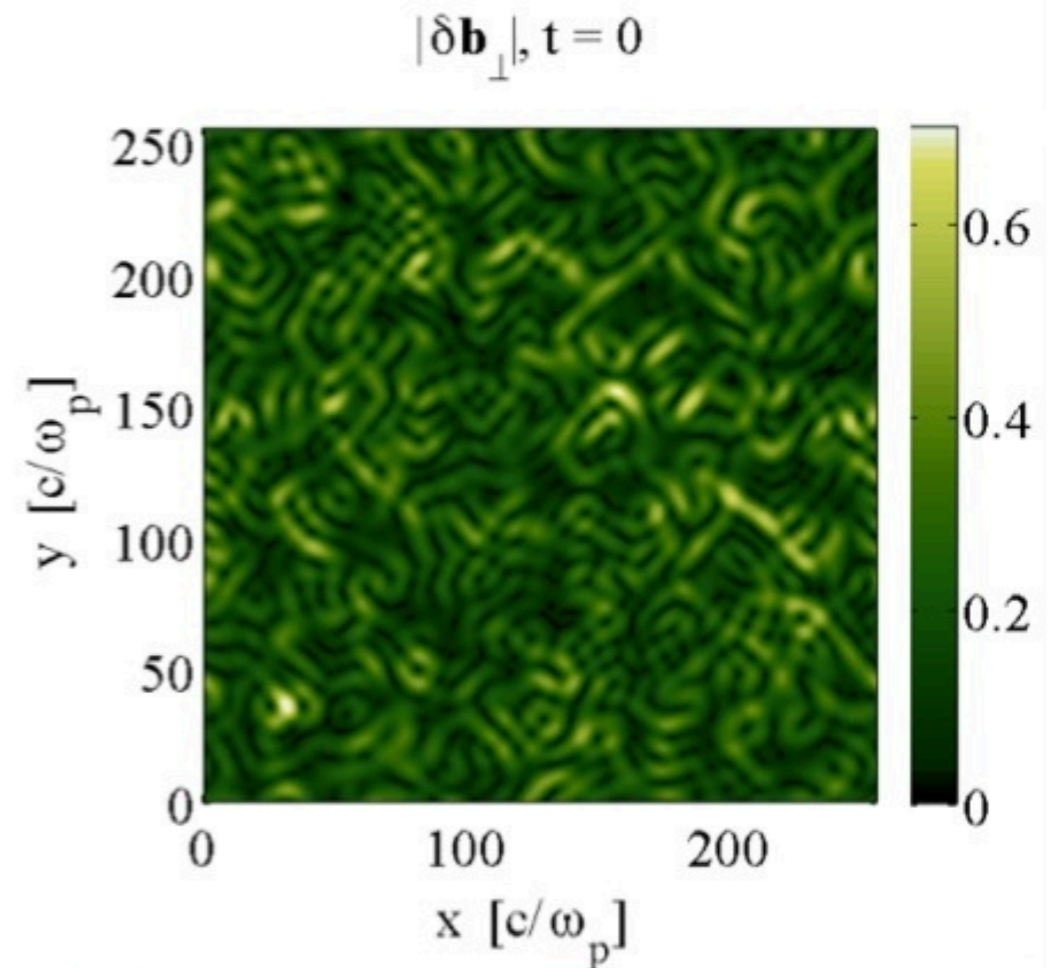
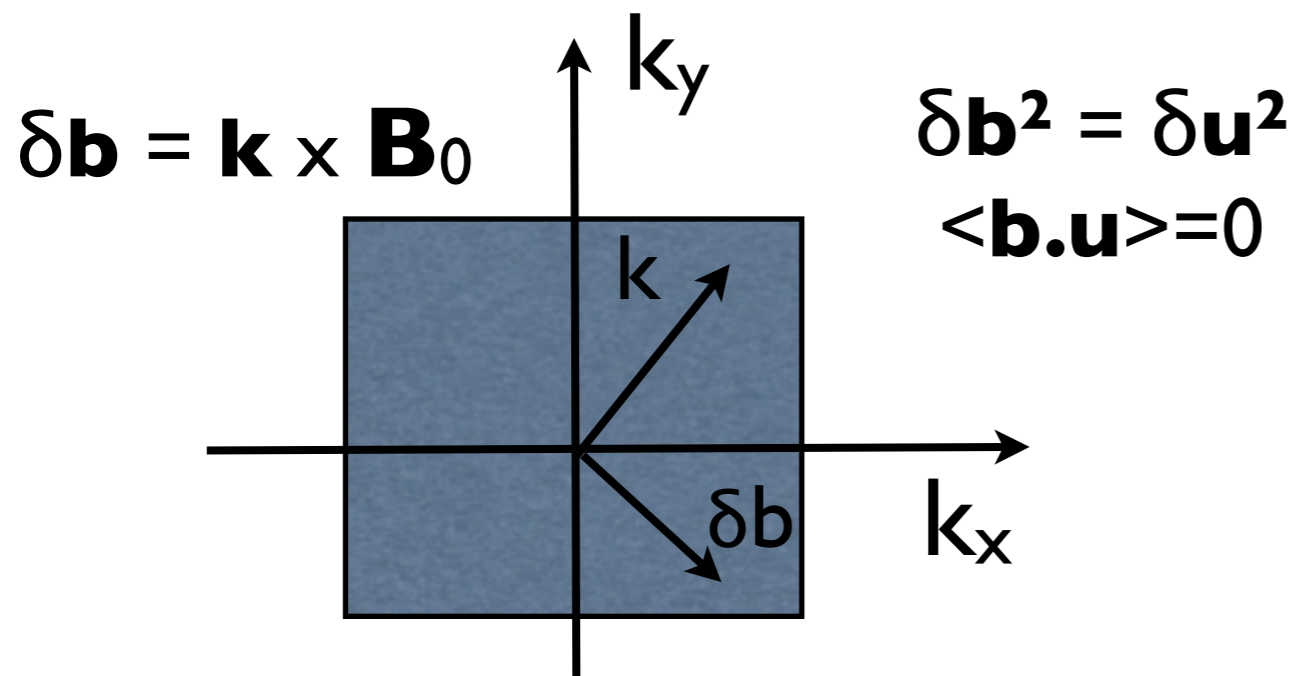
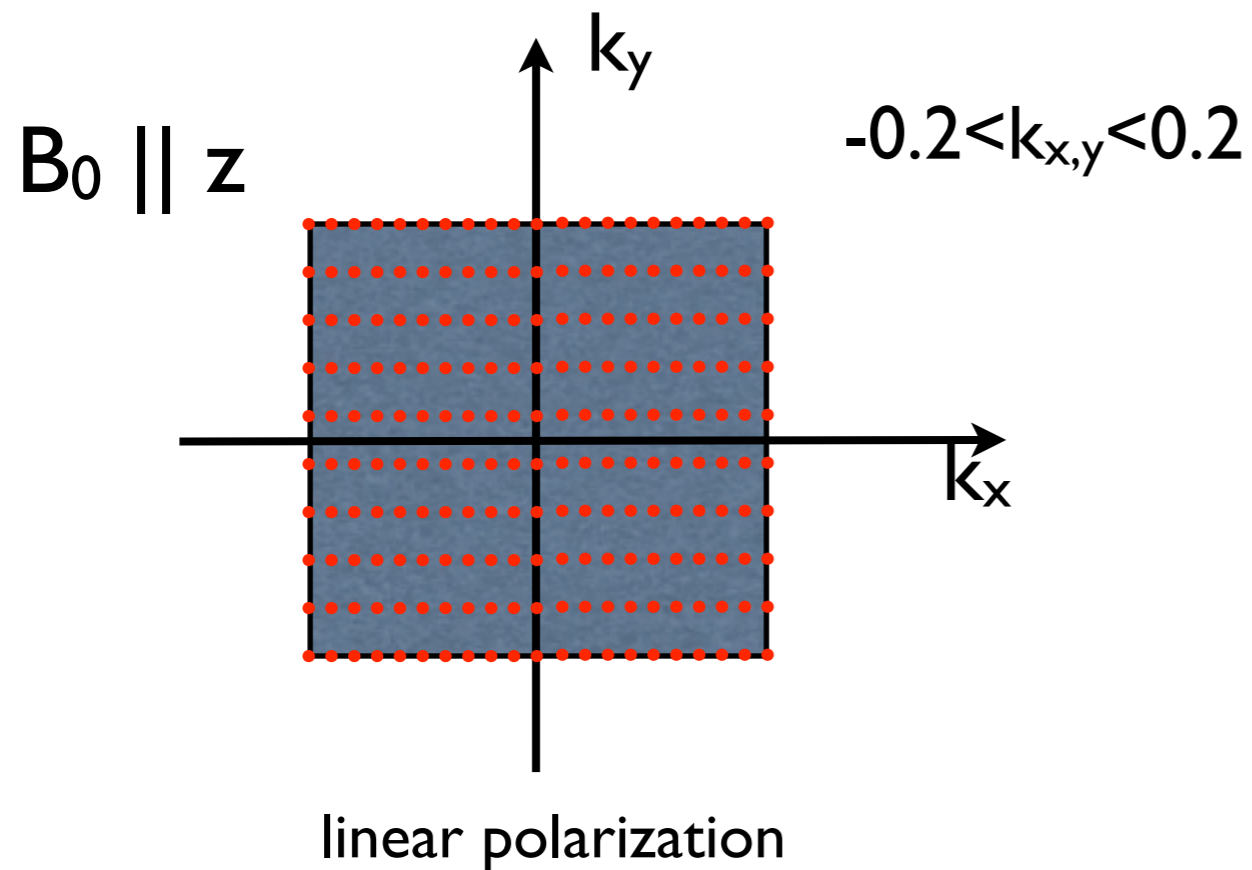
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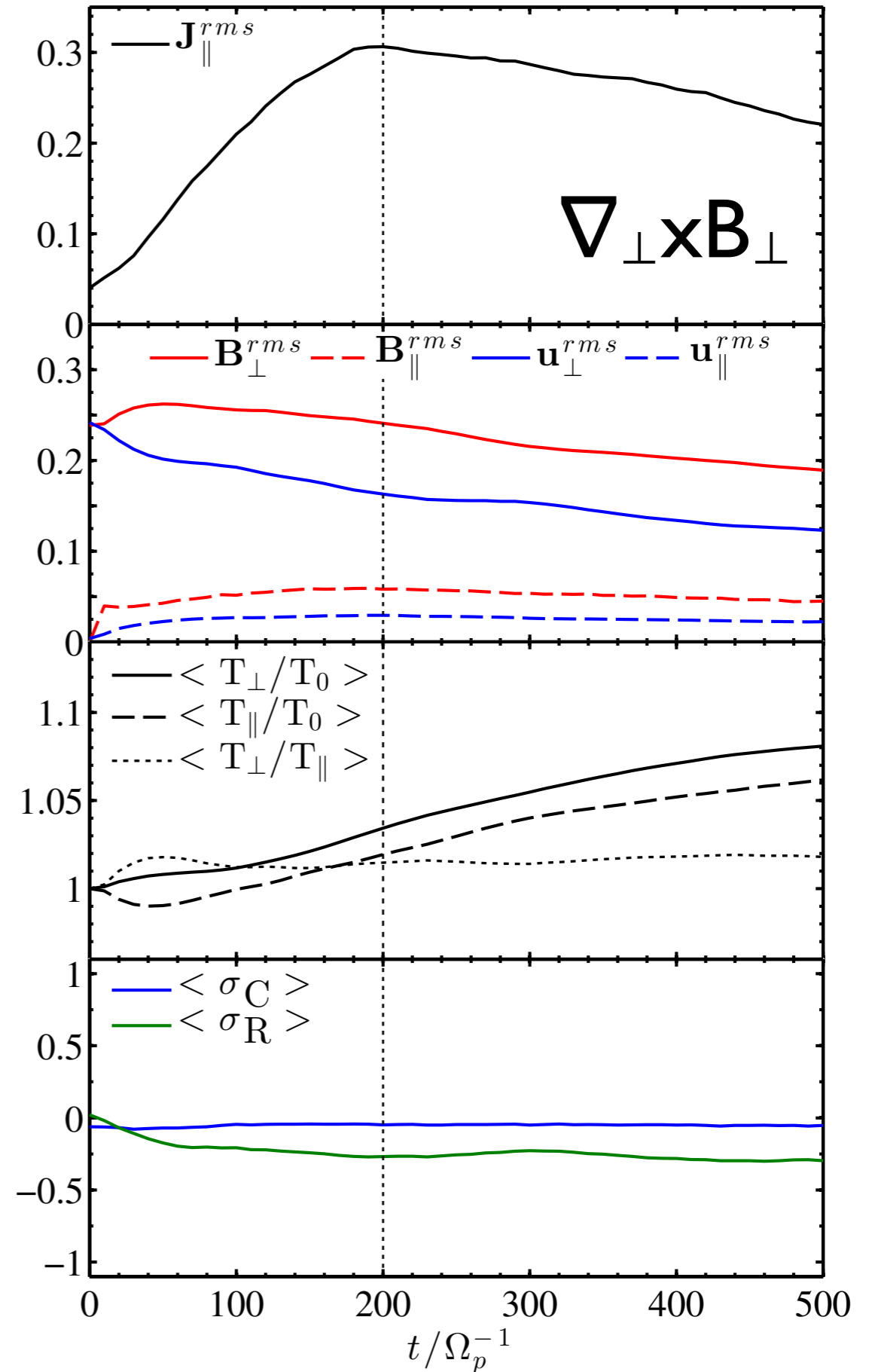
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# Temporal evolution

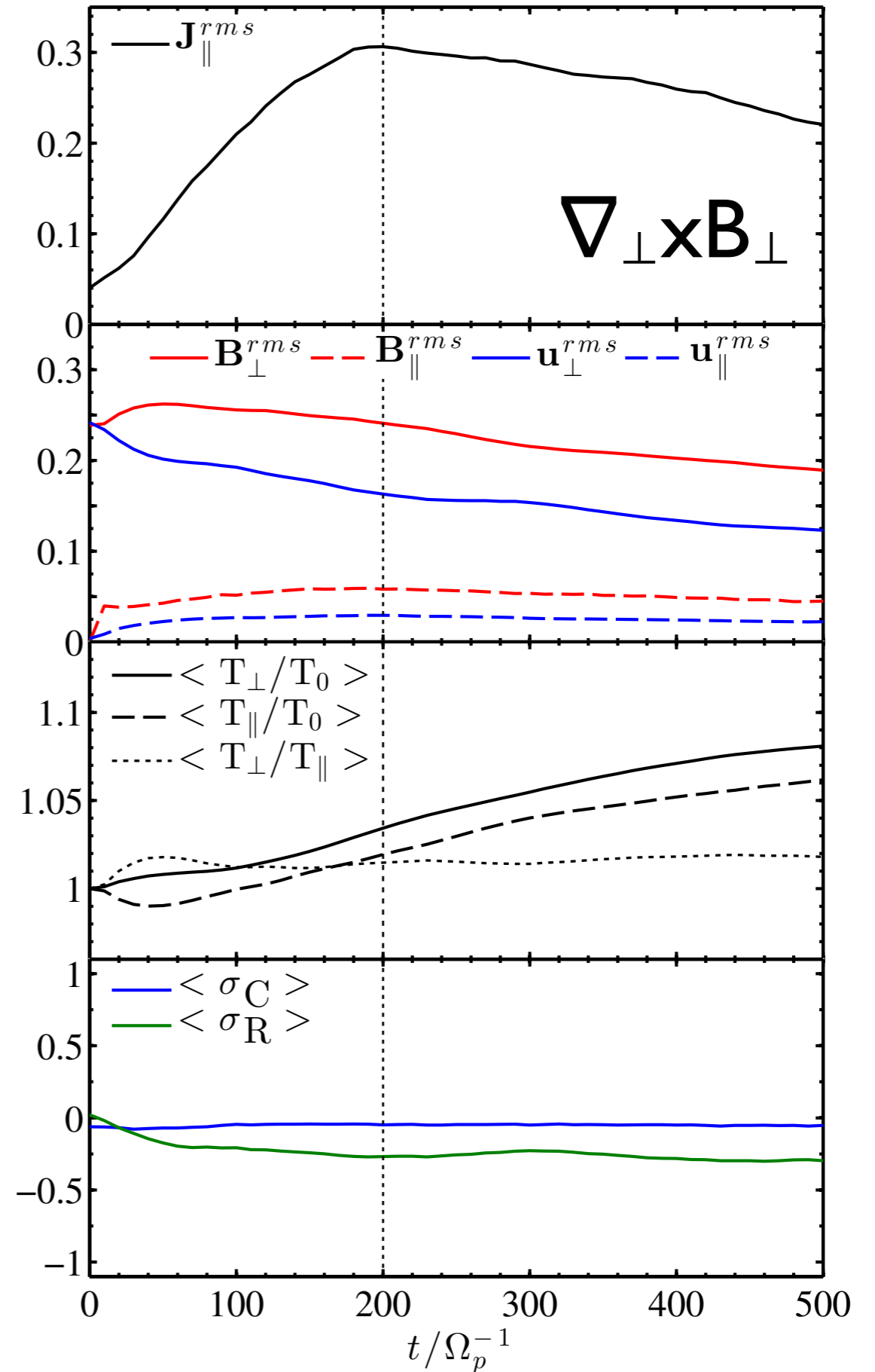
Peak in out-of-plane current identifies  
maximum of turbulent activity  
(*Servidio et al 2012*)



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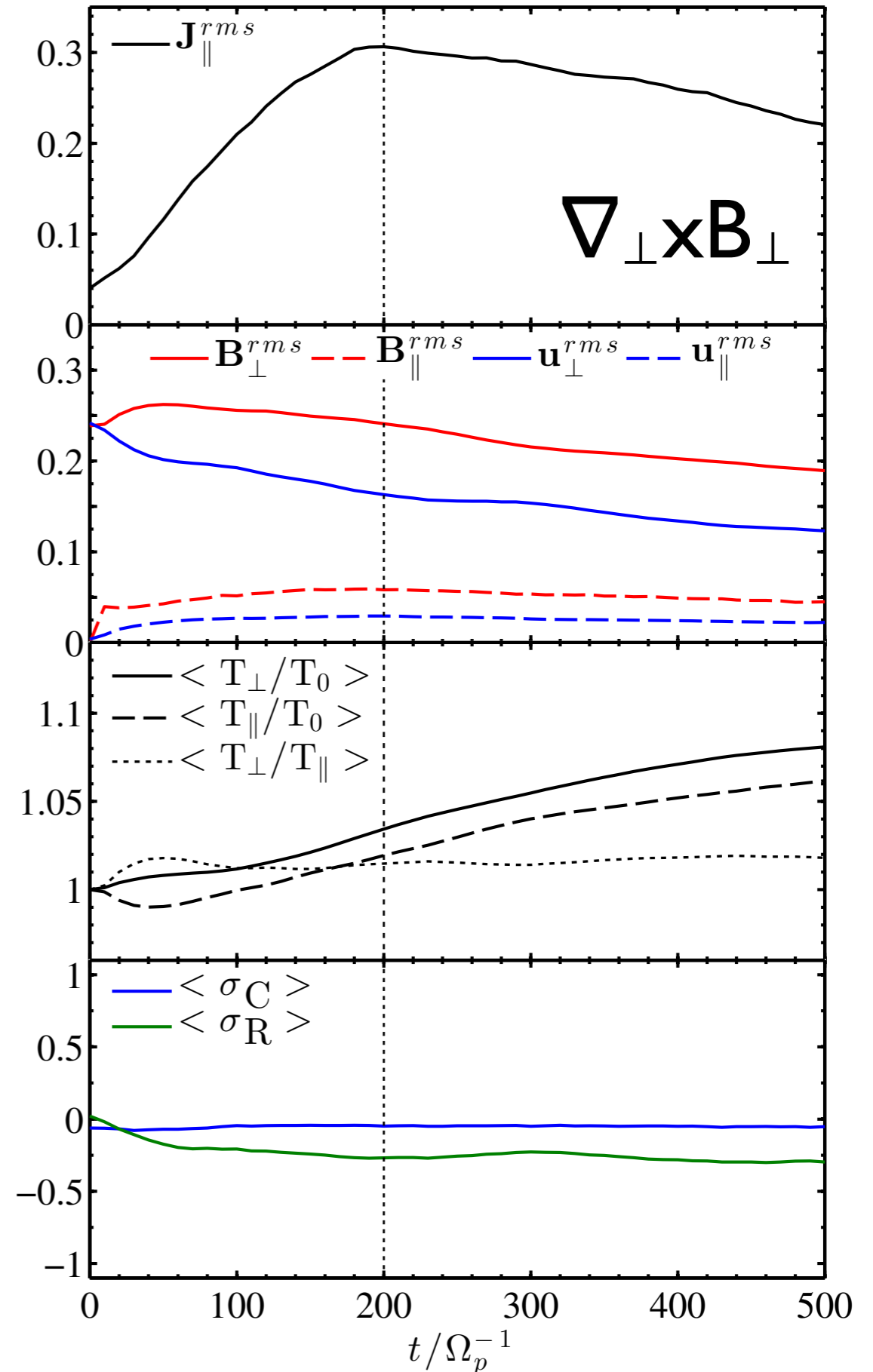


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Anisotropy and parallel/perp temperatures: **heating!**  
but no global anisotropy





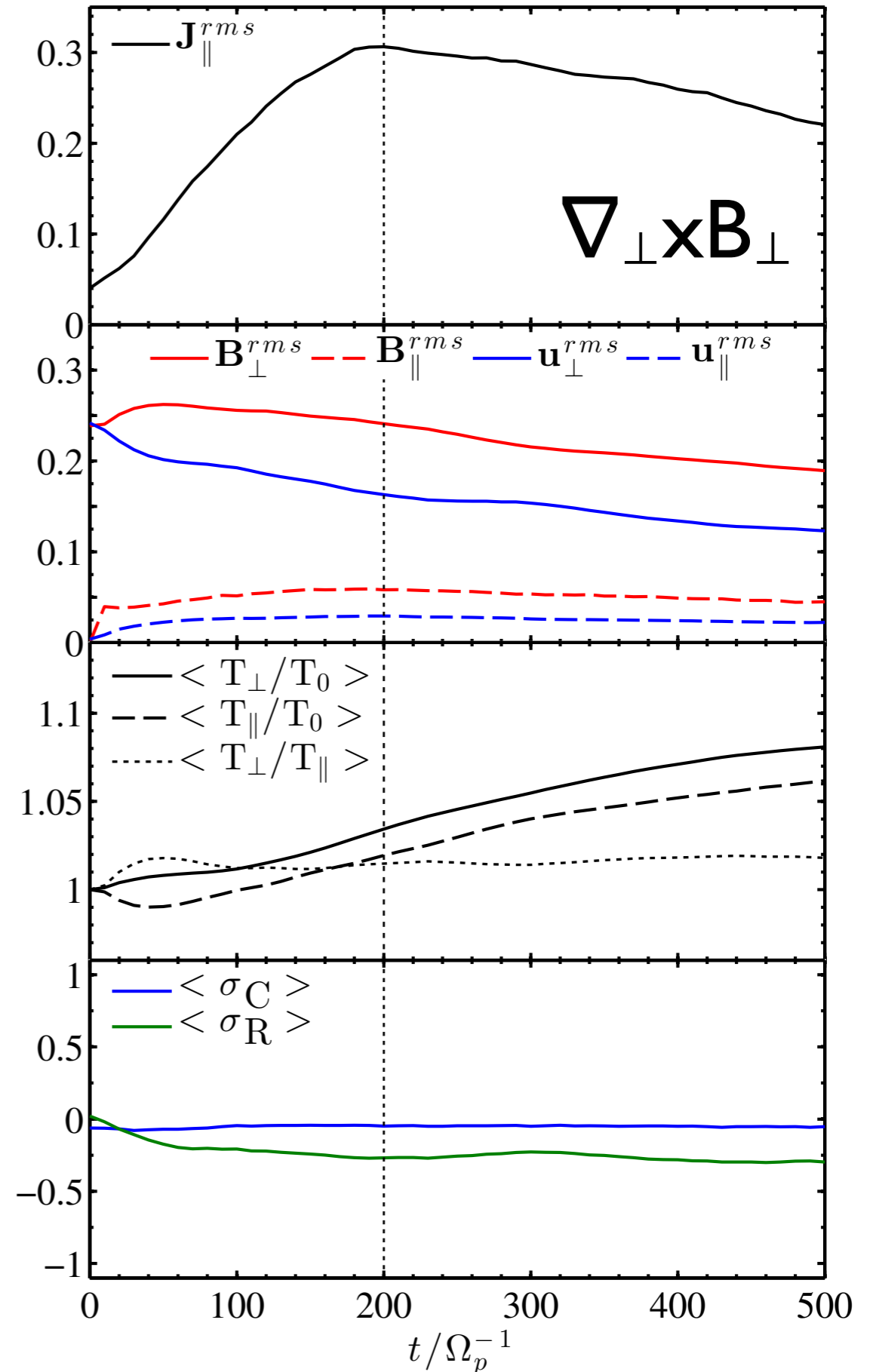
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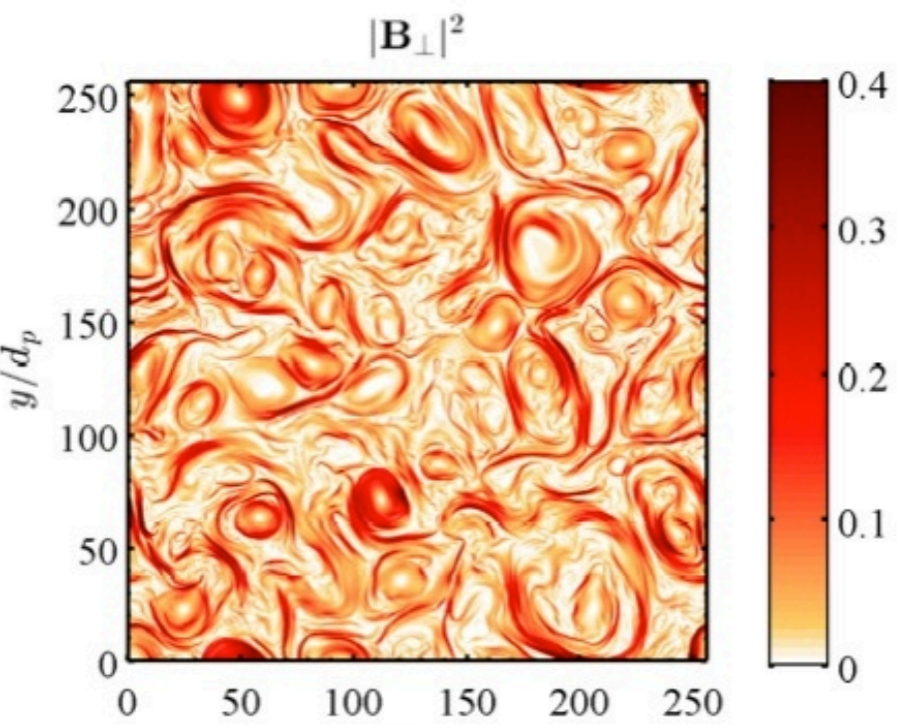
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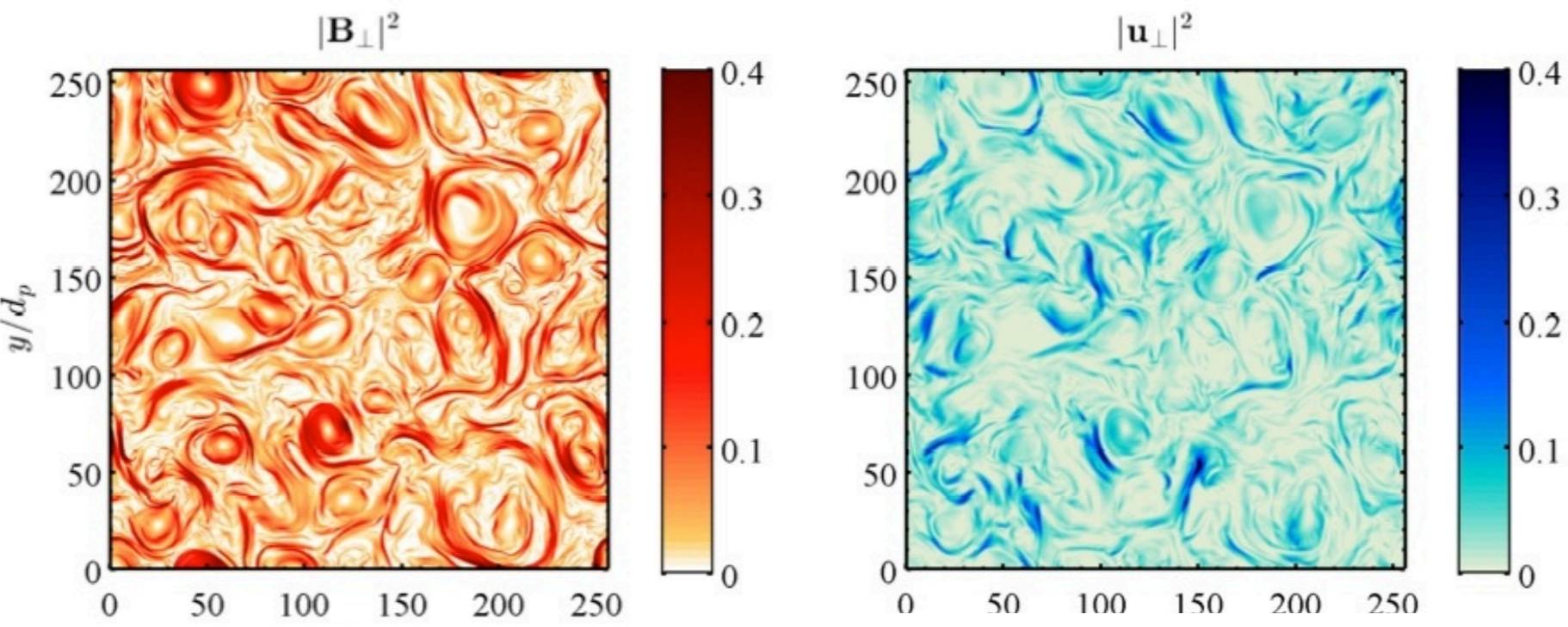
$z^+$  and  $z^-$  remain balanced in time  
while small excess  $b^2 > u^2$



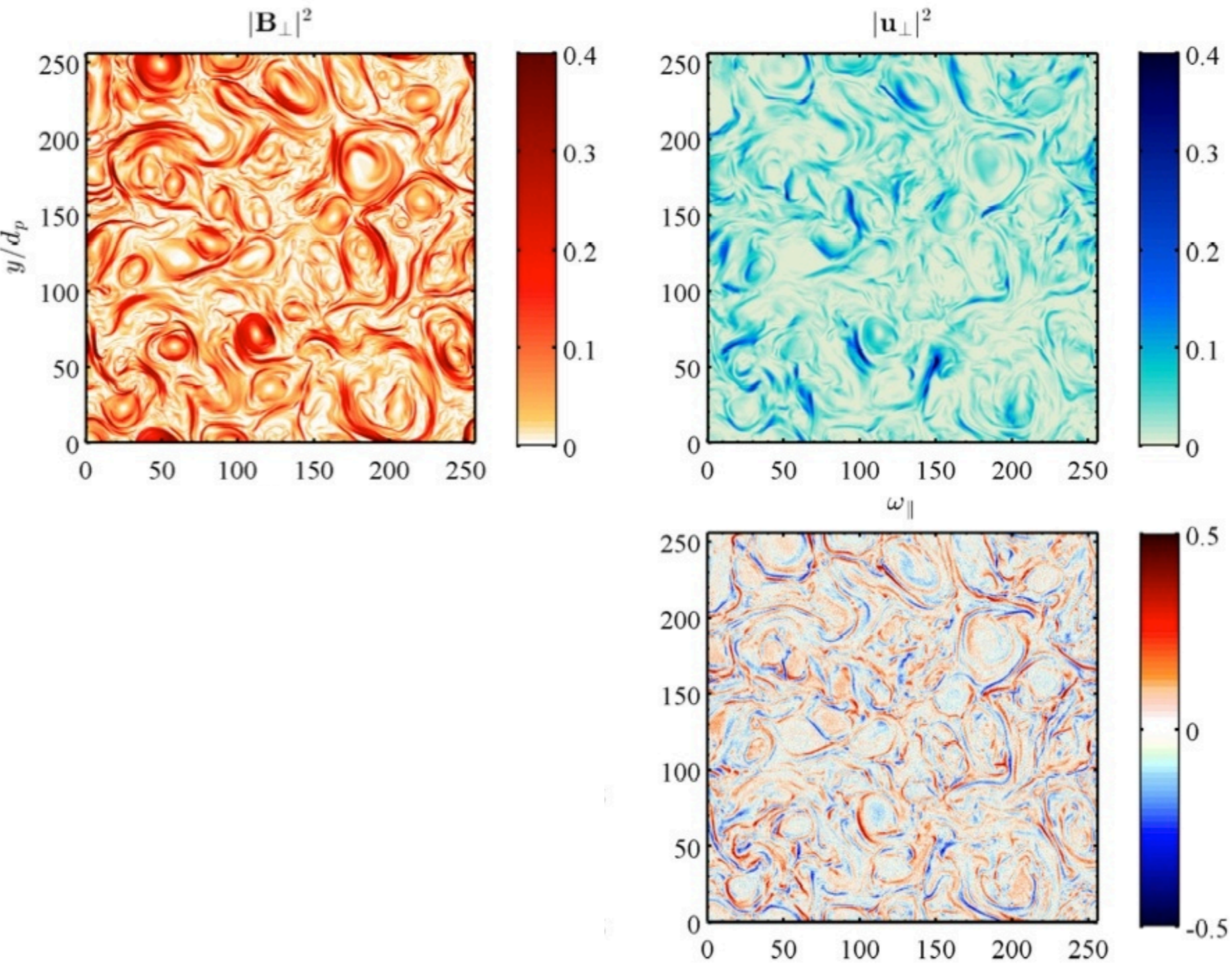
# 2D maps of fields



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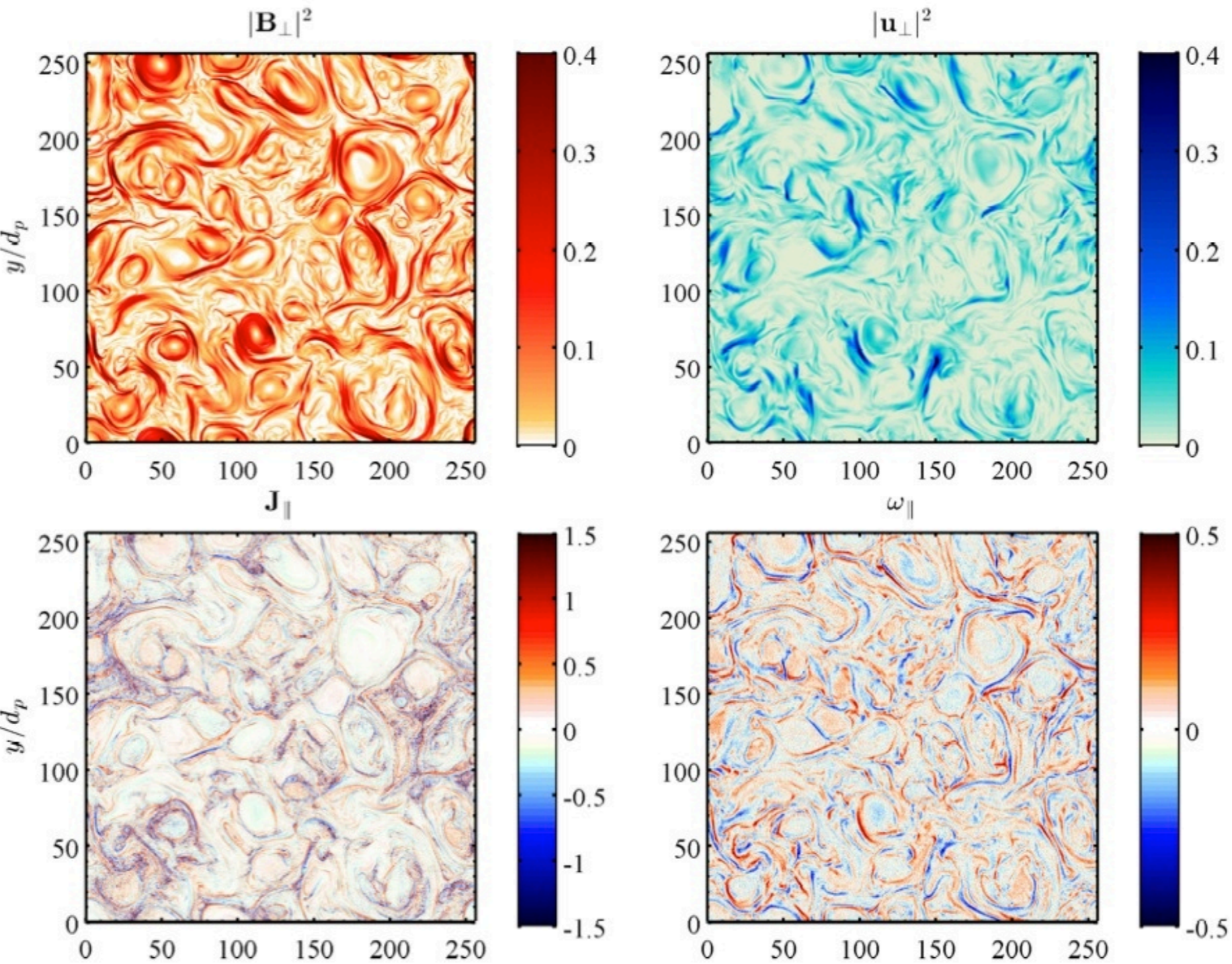


# 2D maps of fields



vorticity  $\nabla \times \mathbf{V}$

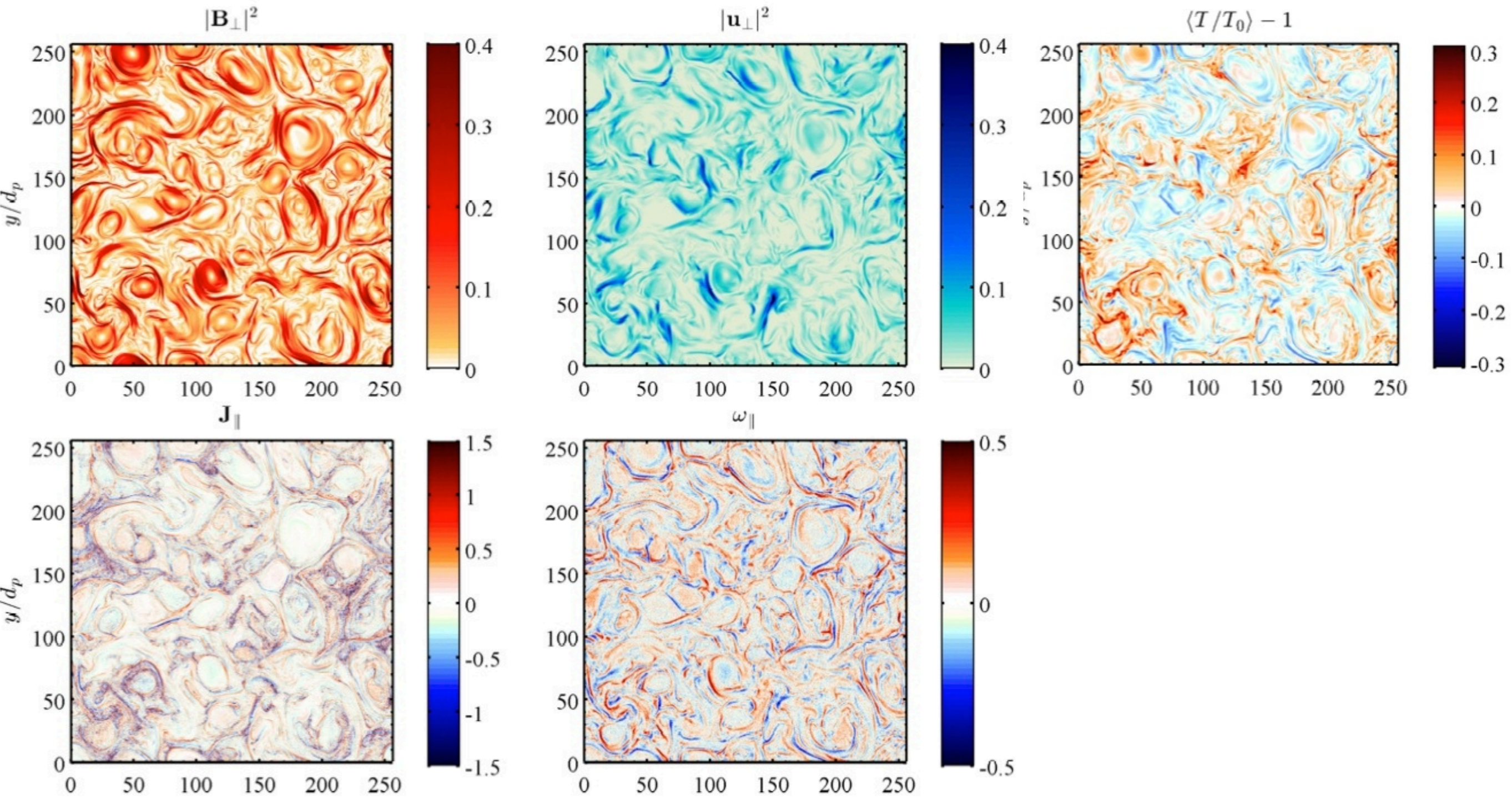
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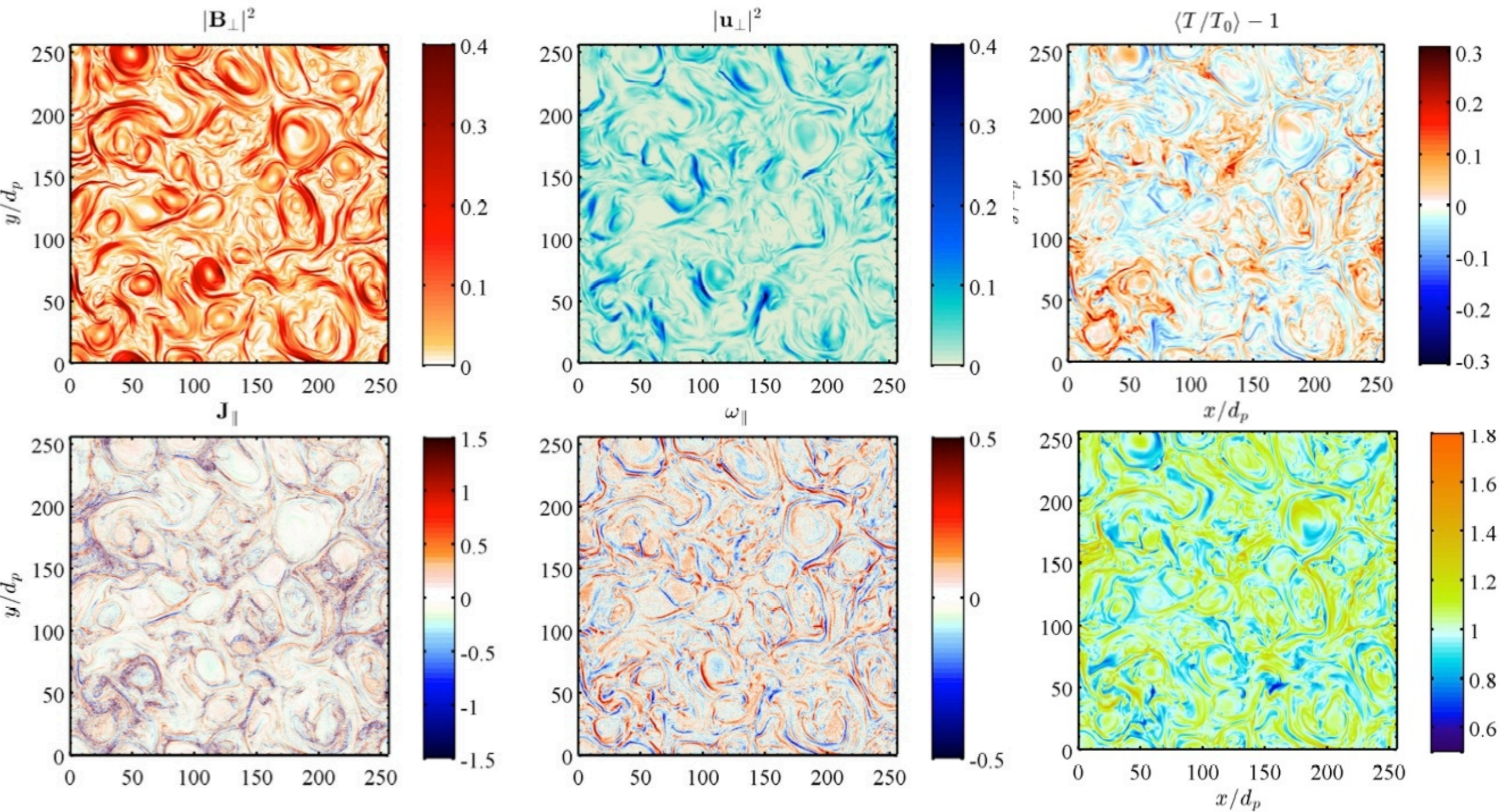
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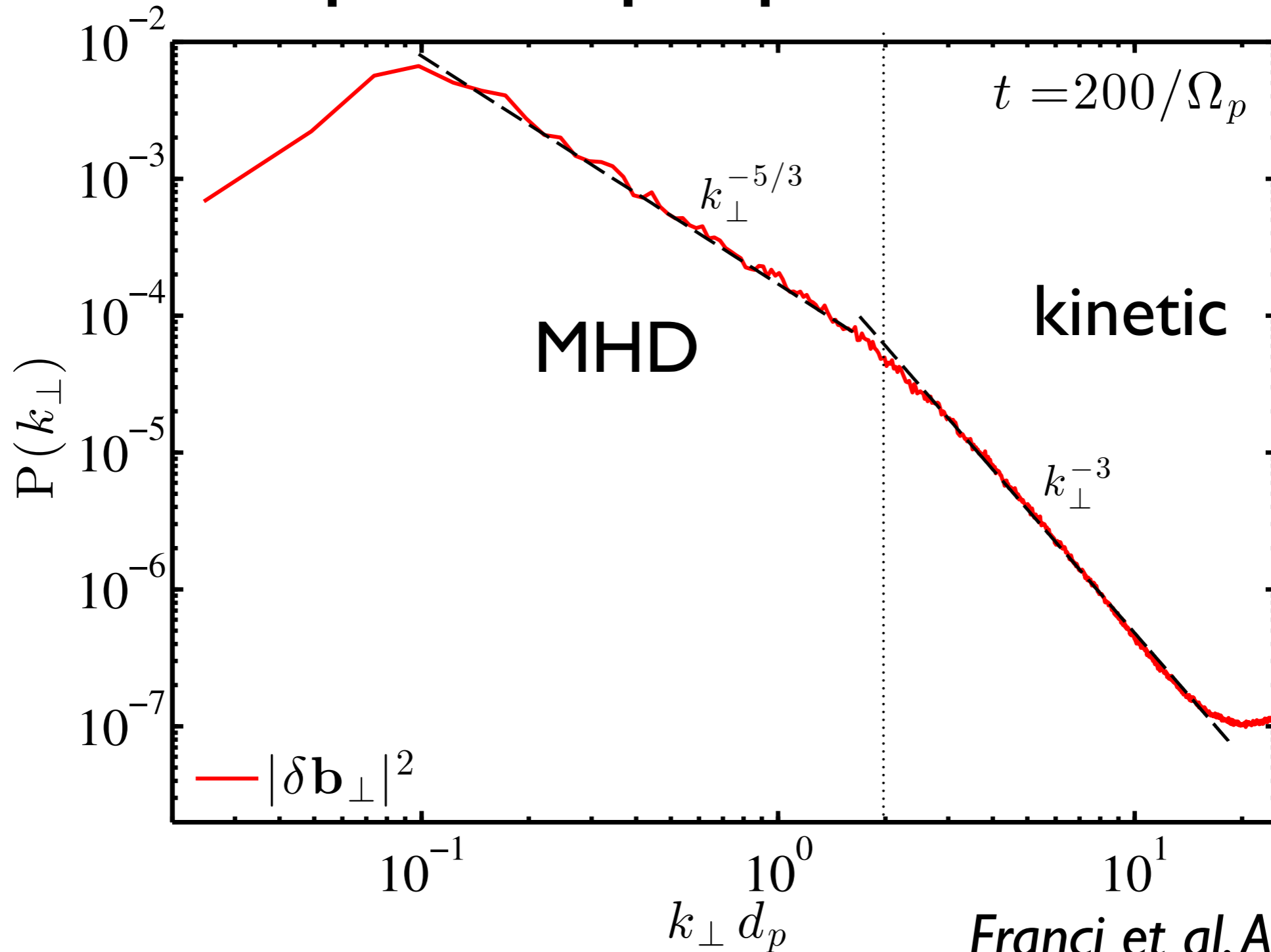


current  $\nabla \times \mathbf{B}$

vorticity  $\nabla \times \mathbf{V}$

Temperature anisotropy  
(Servidio, Perrone, ...)

# Spectral properties

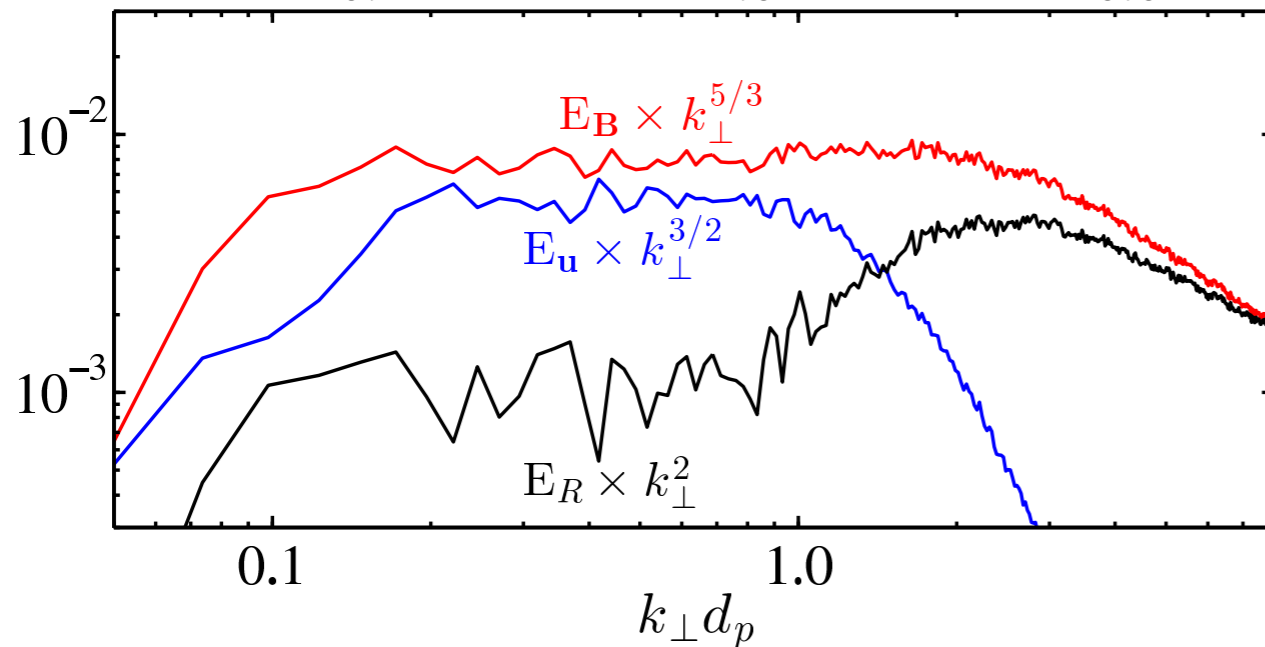
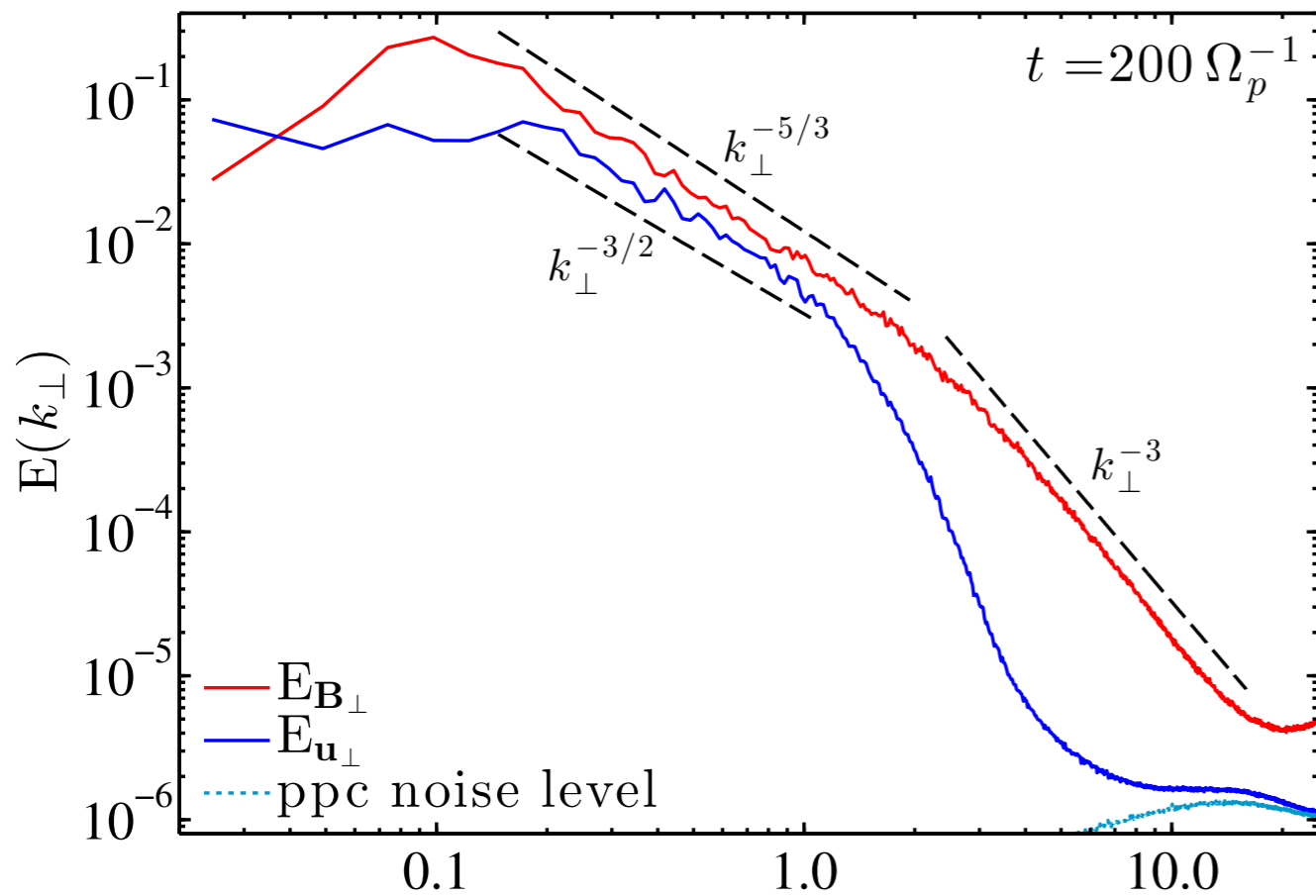


*Franci et al. ApJL 2015*

Self-consistent transition between two turbulent regimes:  
Kolmogorov-like at large MHD scales and kinetic at sub-ion scales



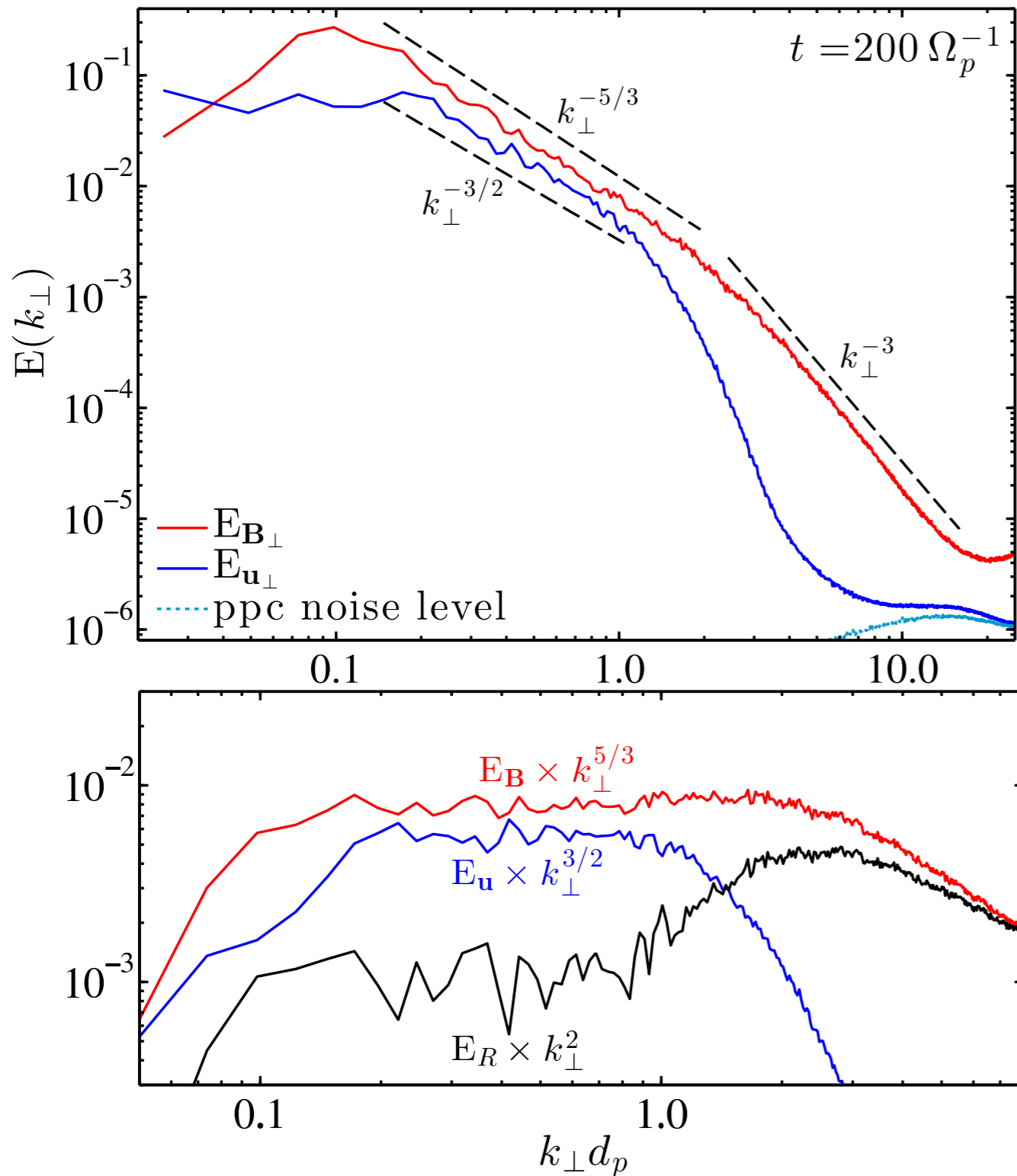
# Spectrum of V and residual energy



**-5/3, -3/2, -2**

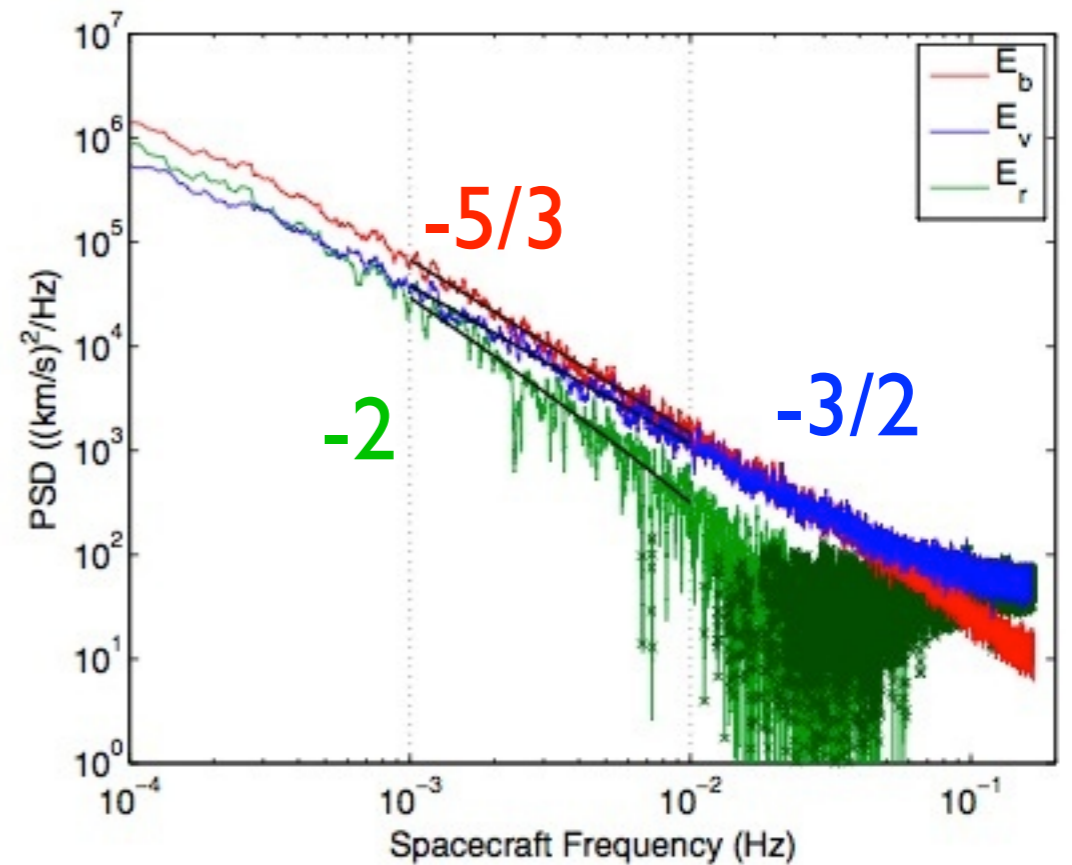
Different scaling for B and V in the MHD inertial range. Their difference (**residual energy**) is also a power law! (Roland's talk)

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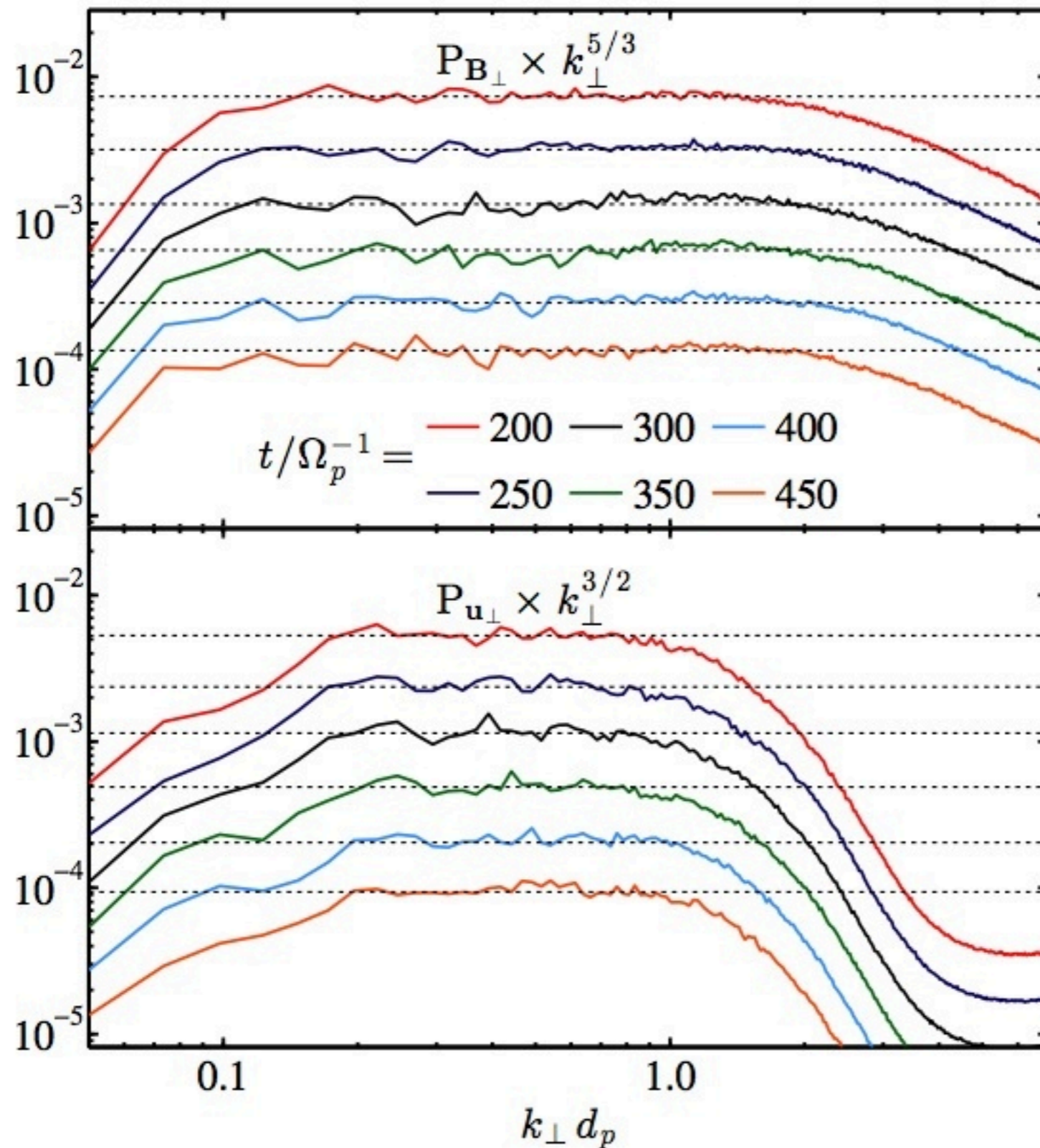
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*Chen et al. 2013*

# Stability of the turbulent cascade

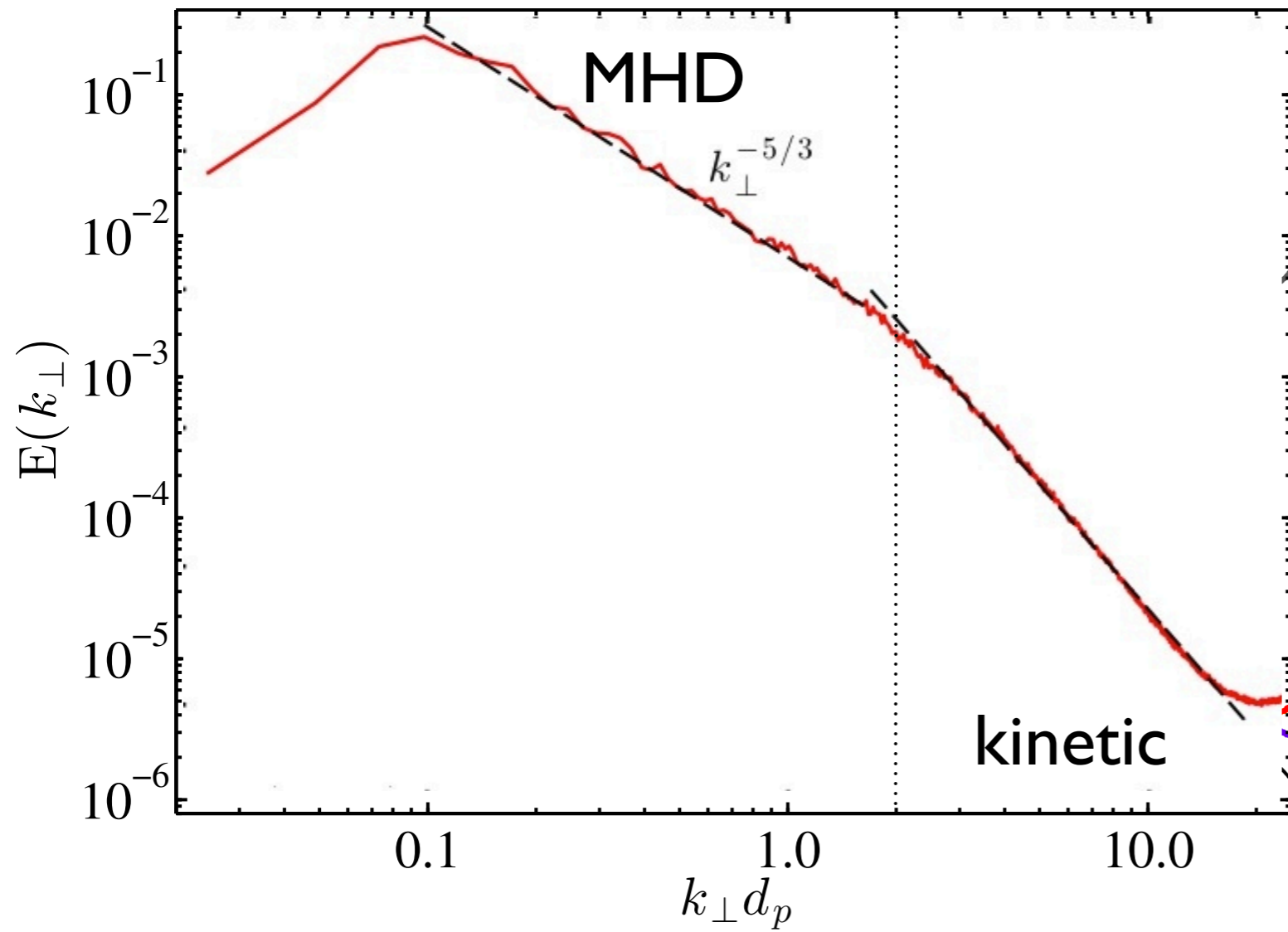
Compensated spectra at different simulation times



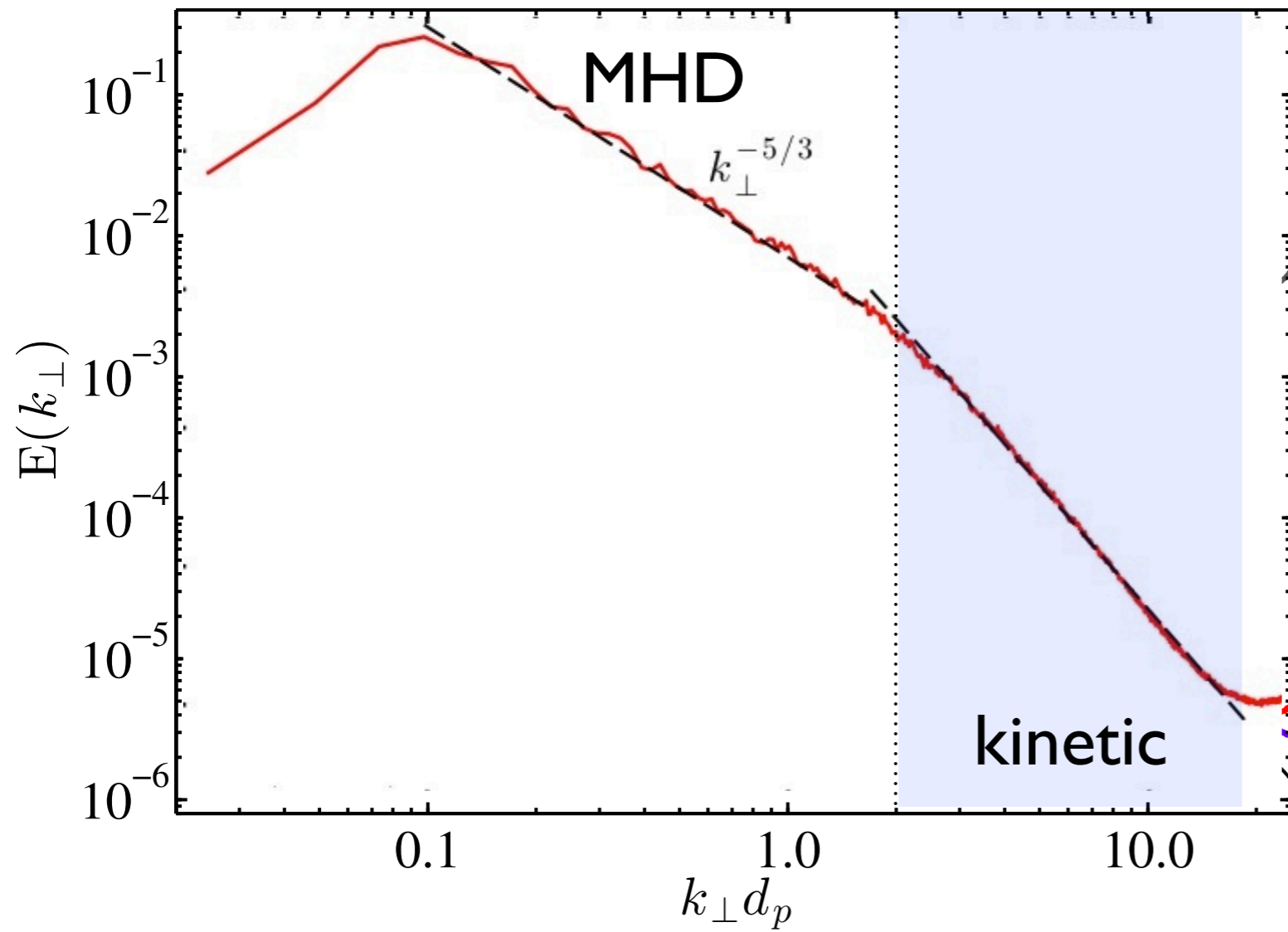
Magnetic field  
compensated by 5/3

Proton velocity  
compensated by 3/2

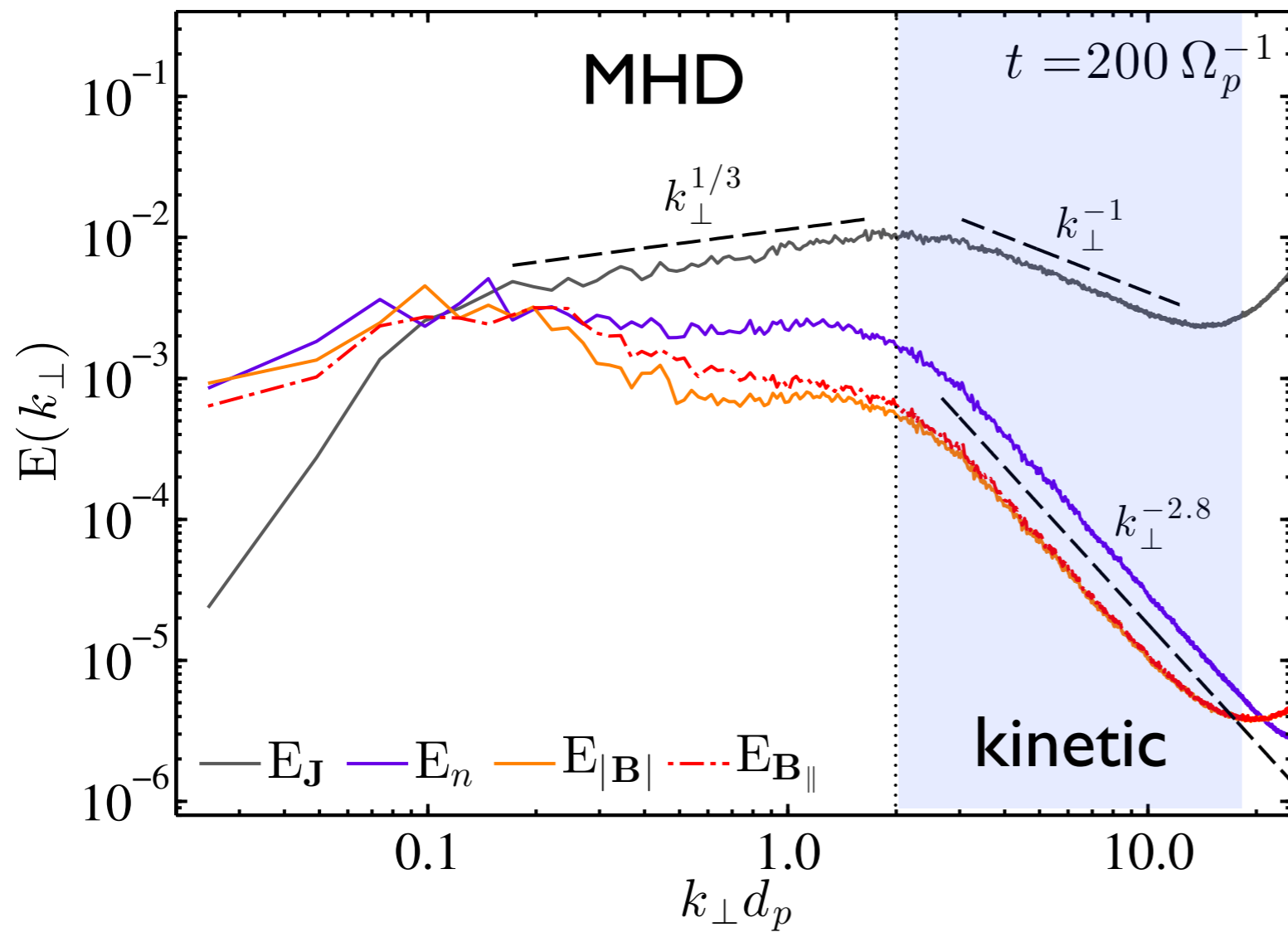
# Turbulence at kinetic scales



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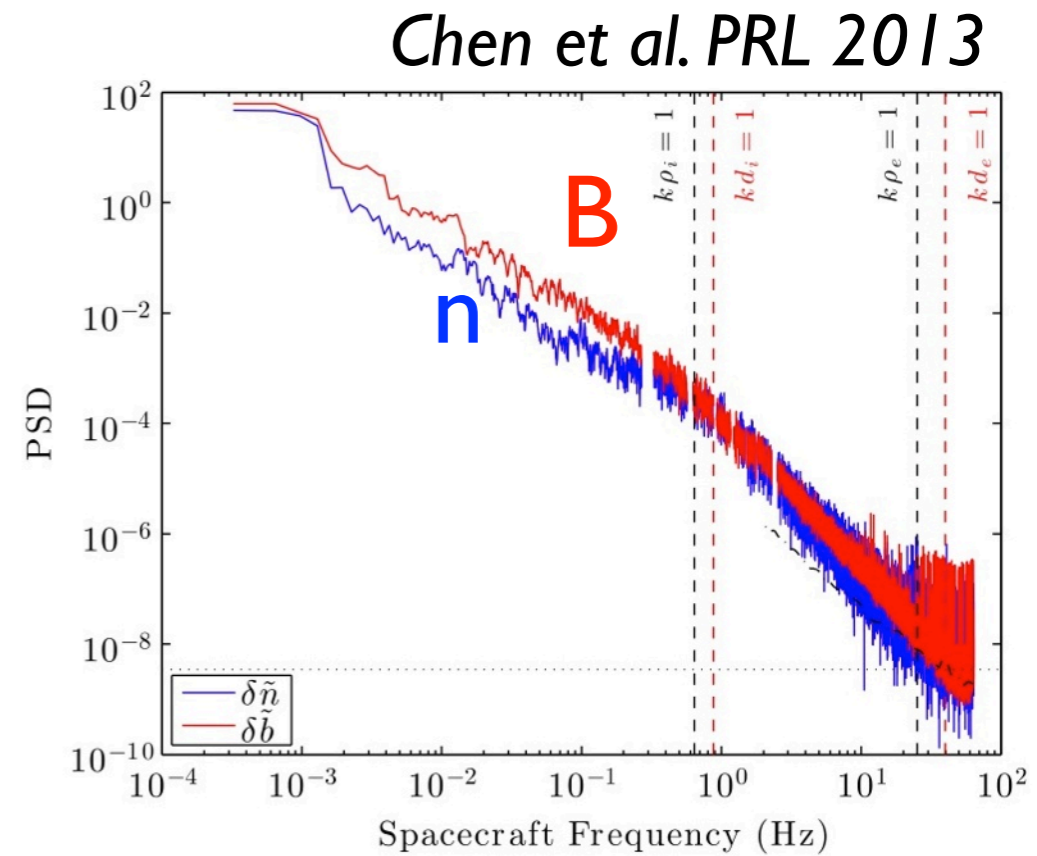
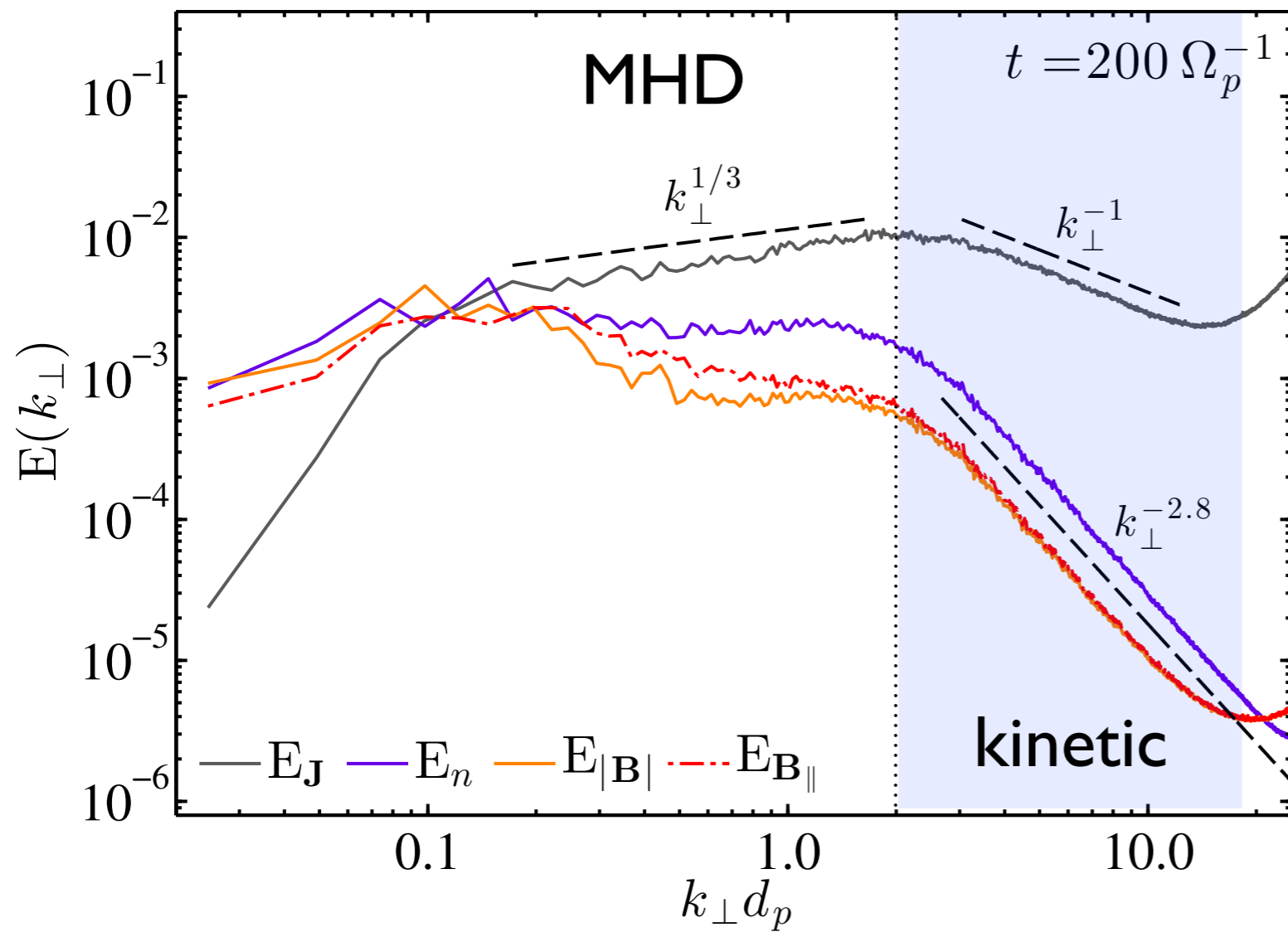
- Current has a peak:

$$\mathbf{J} = \nabla \times \mathbf{B} \sim k \delta \mathbf{b} \rightarrow \delta J^2 \sim k^2 \delta b^2$$

- $\delta b_{\parallel} \sim \delta n \propto k^{-2.8}$

- $\delta b_{\parallel} \sim \delta b_{\perp}$

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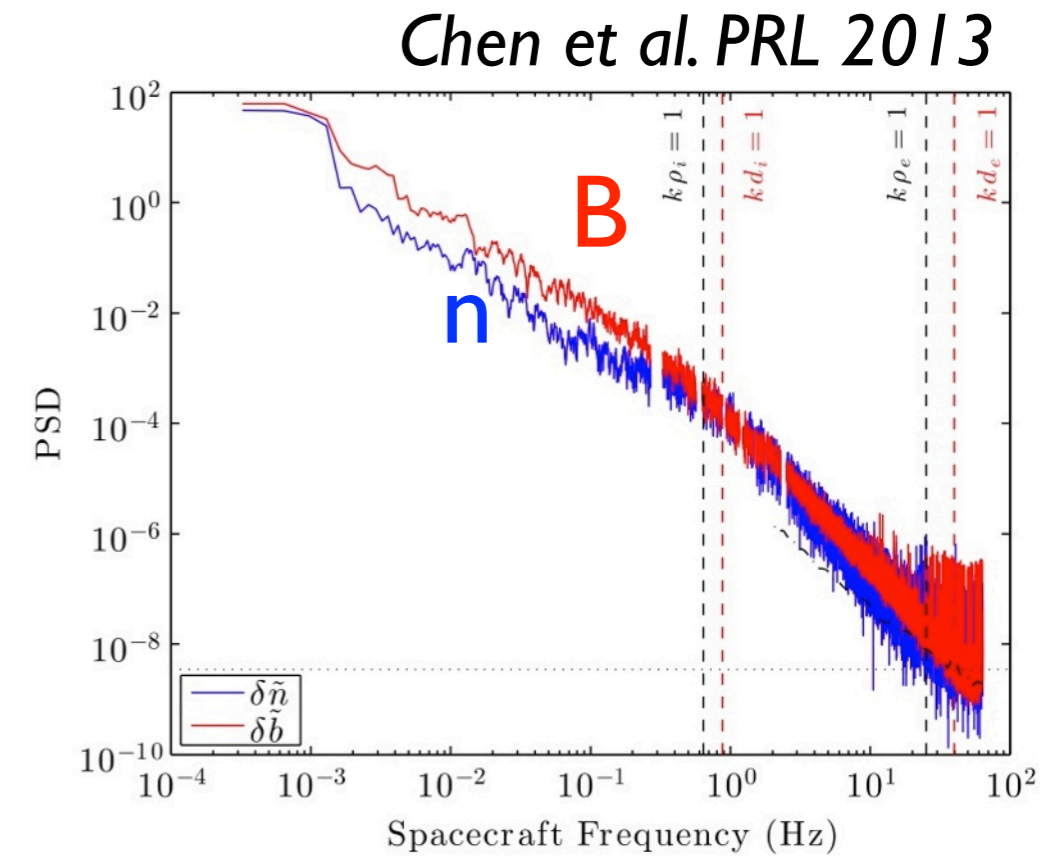
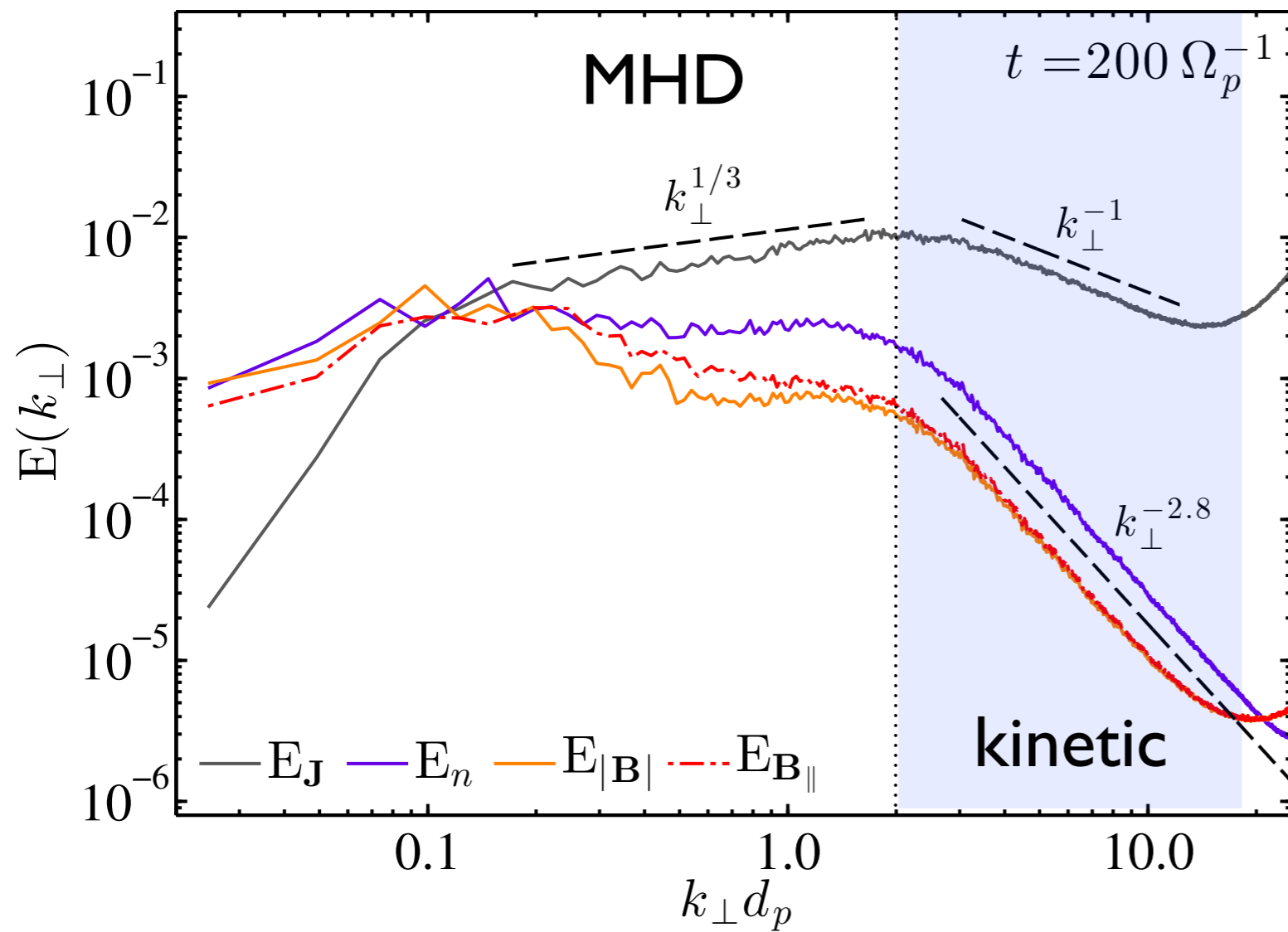
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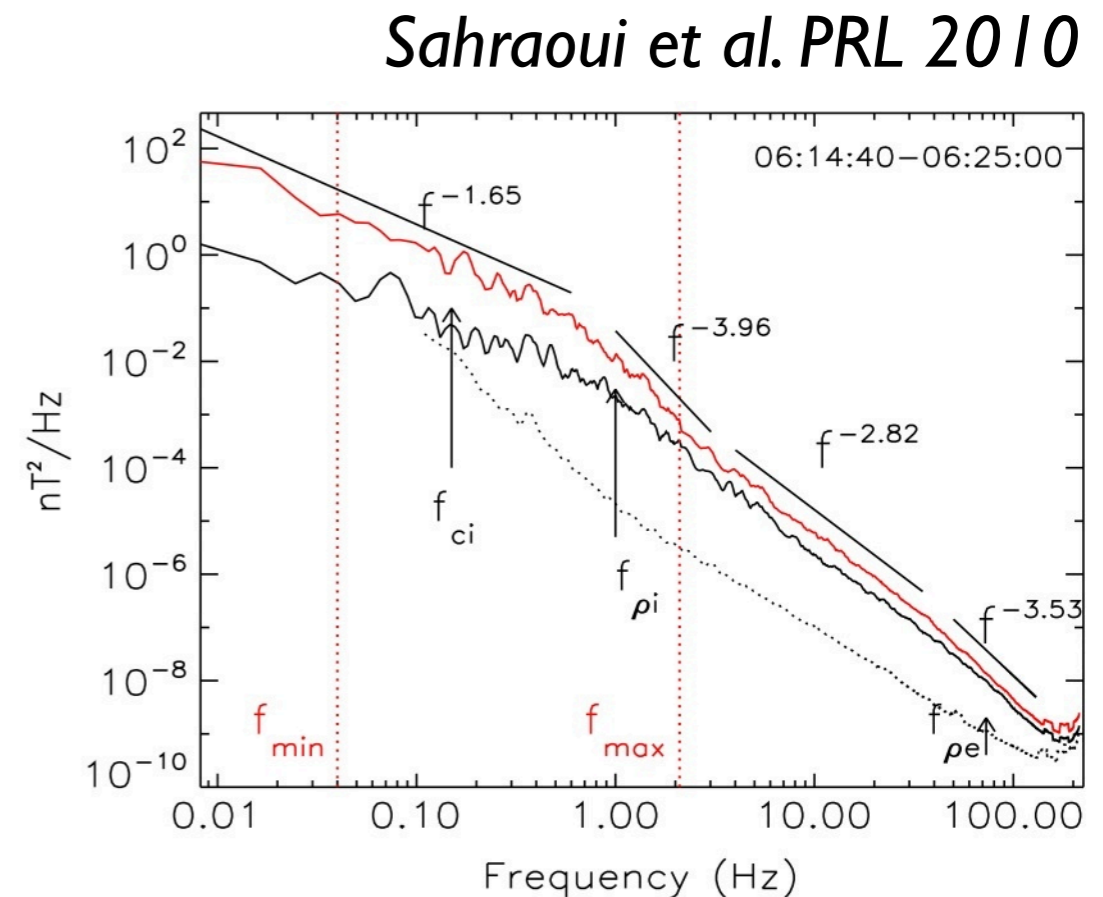


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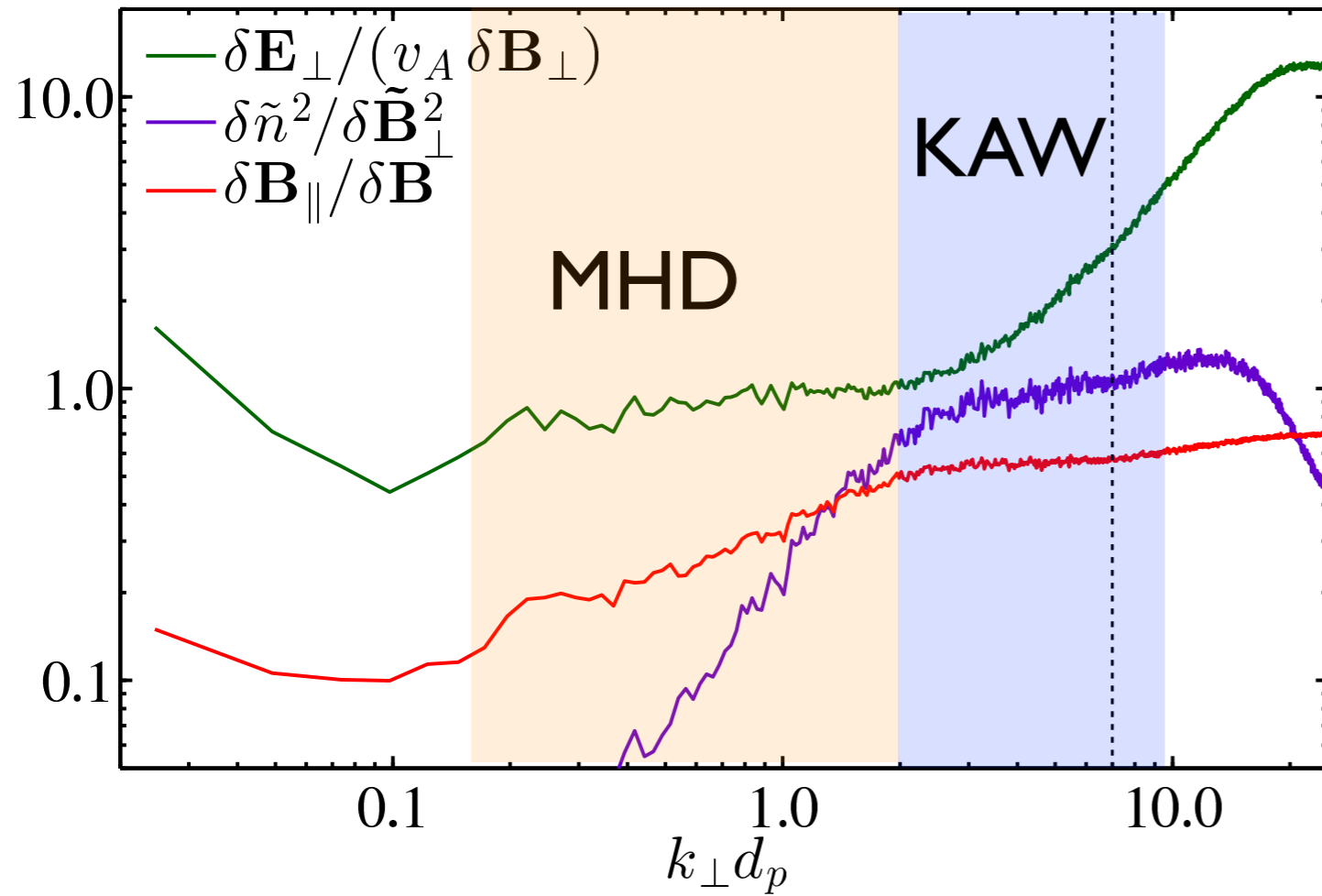
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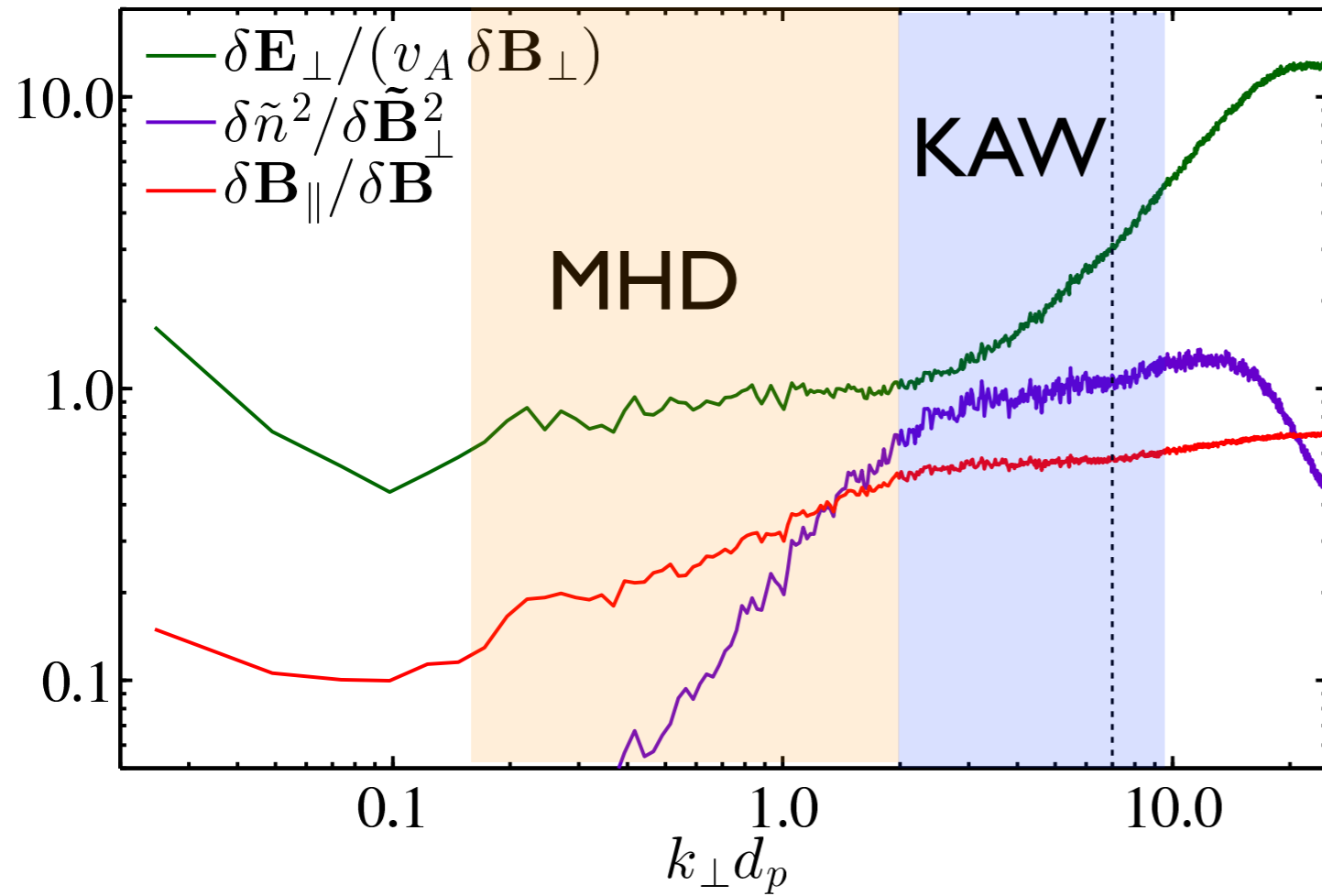


# Nature of sub-ion scale fluctuations?

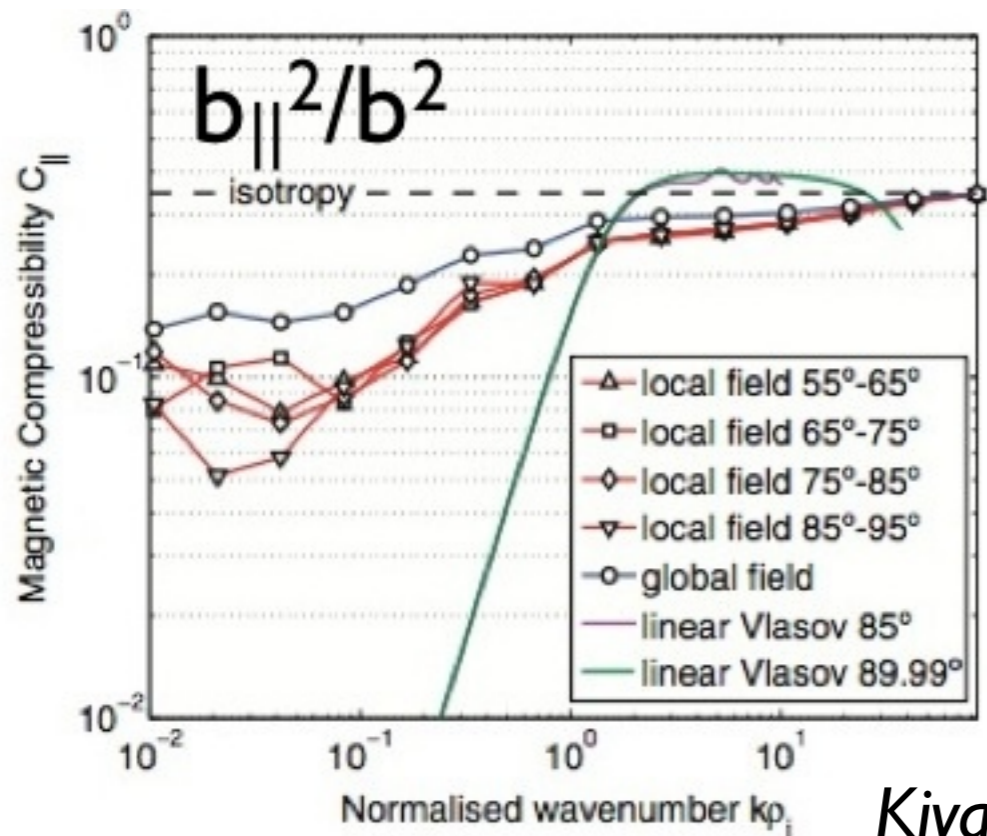


predictions from low frequency kinetic Alfvén wave (KAW) turbulence are very well verified! (which depend on  $\beta_p$  and  $T_p/T_e$ )

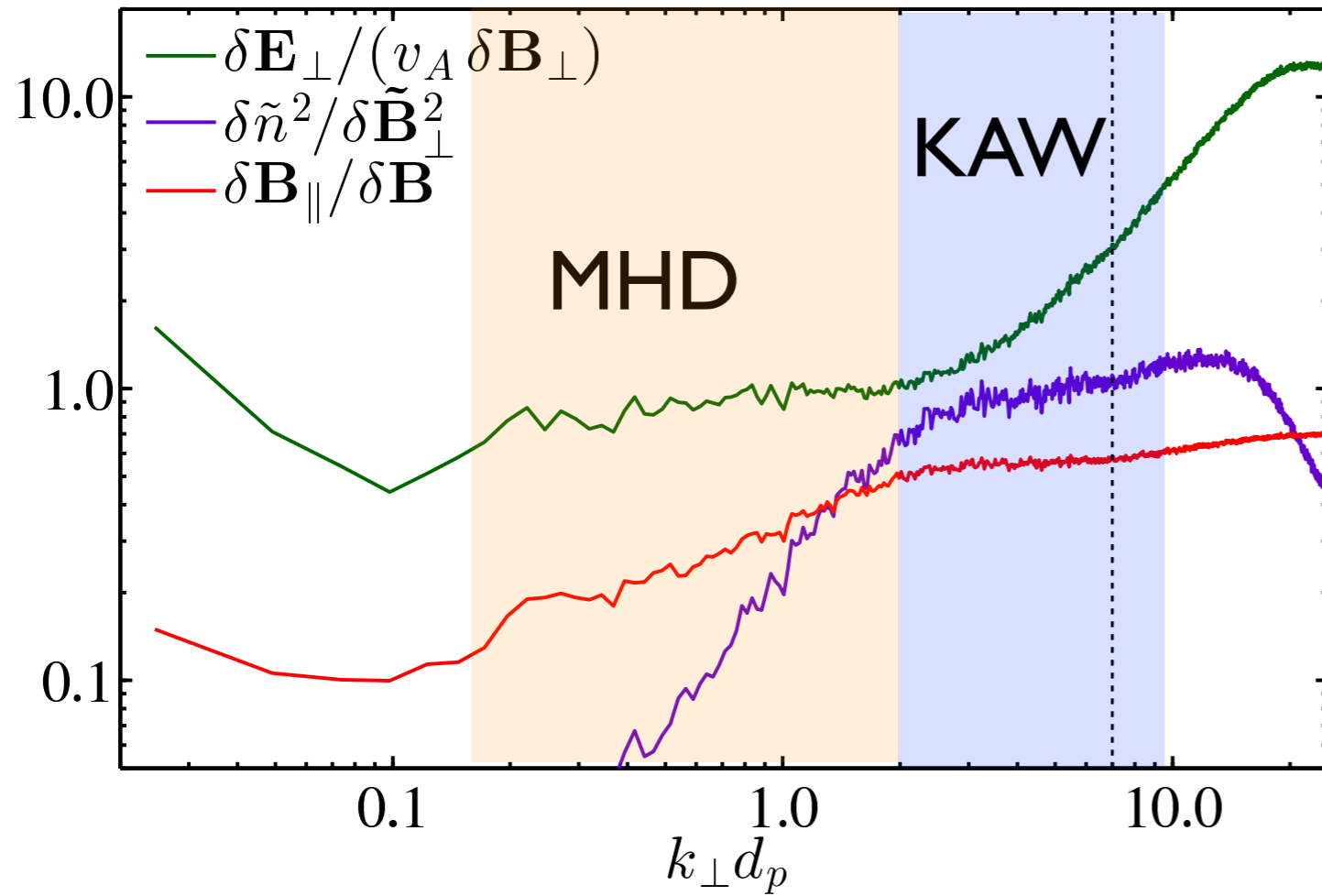
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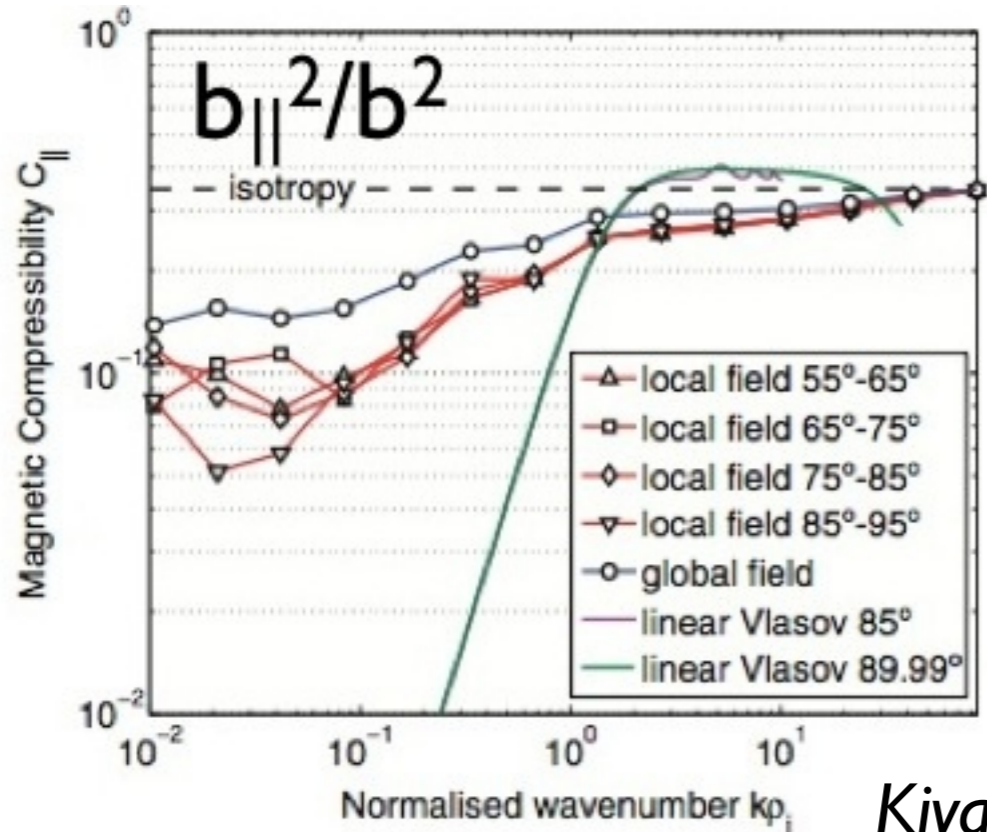
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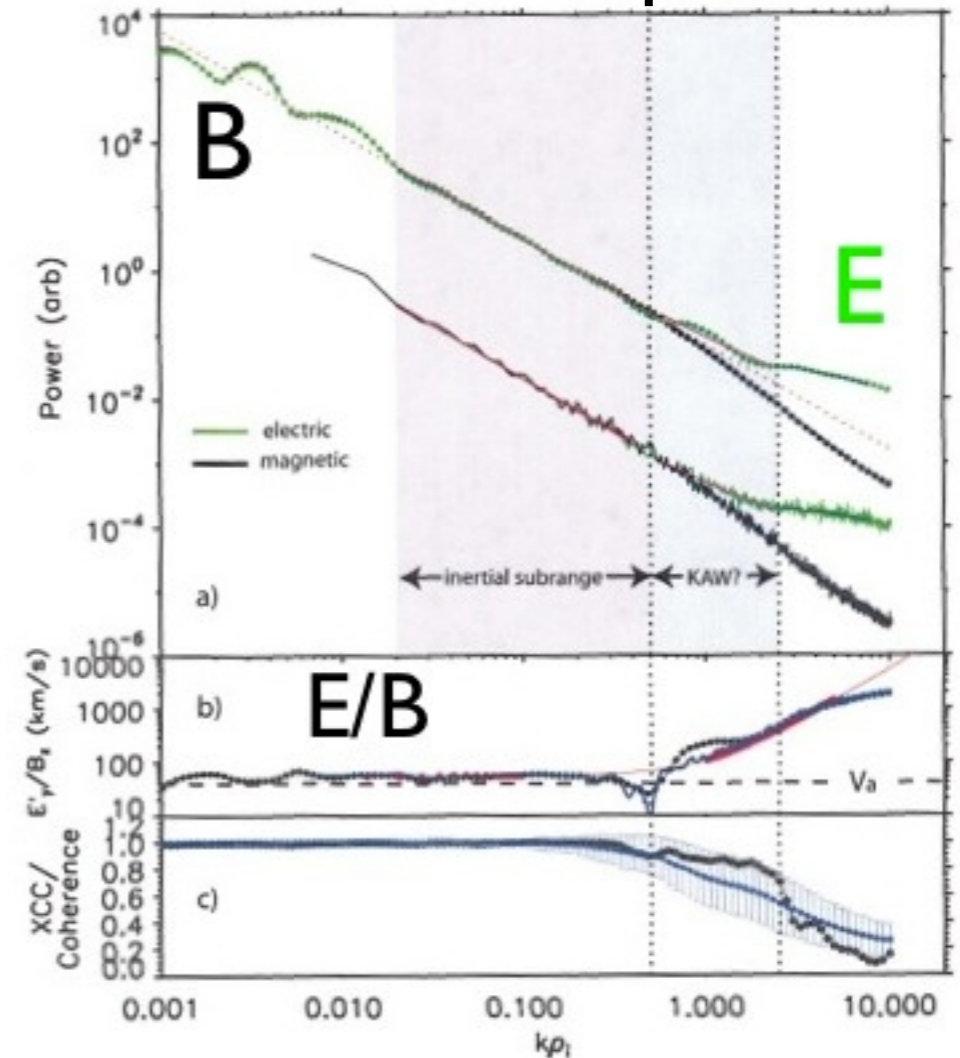


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Kiyani et al. *ApJ* 2013

electric field spectrum



Bale et al *PRL* 2005

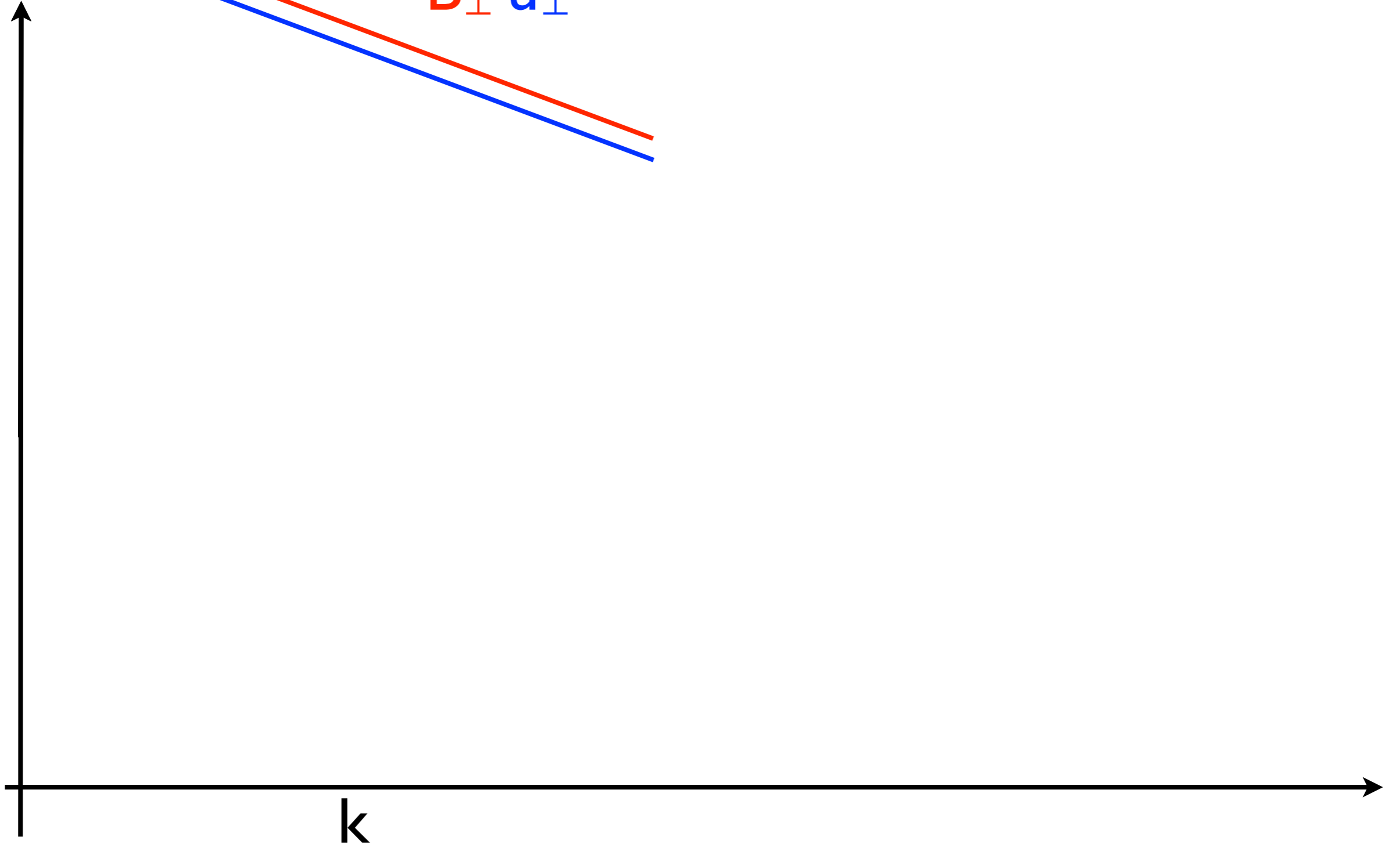
# MHD-kinetic turbulent cascade

Energy



MHD:

$B_{\perp}$   $u_{\perp}$



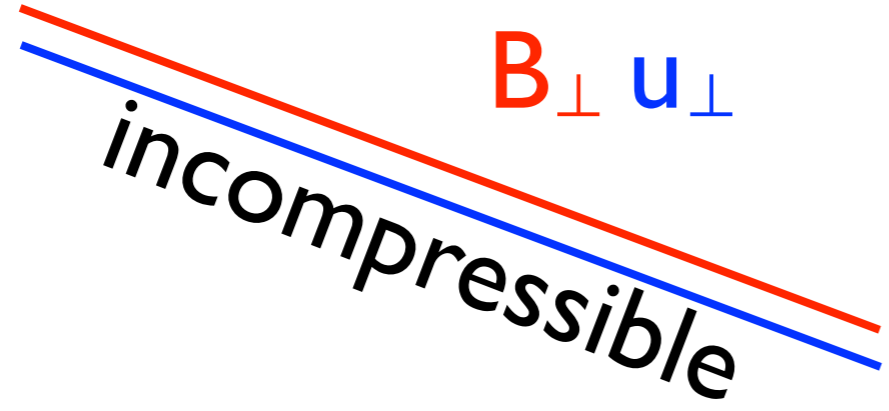
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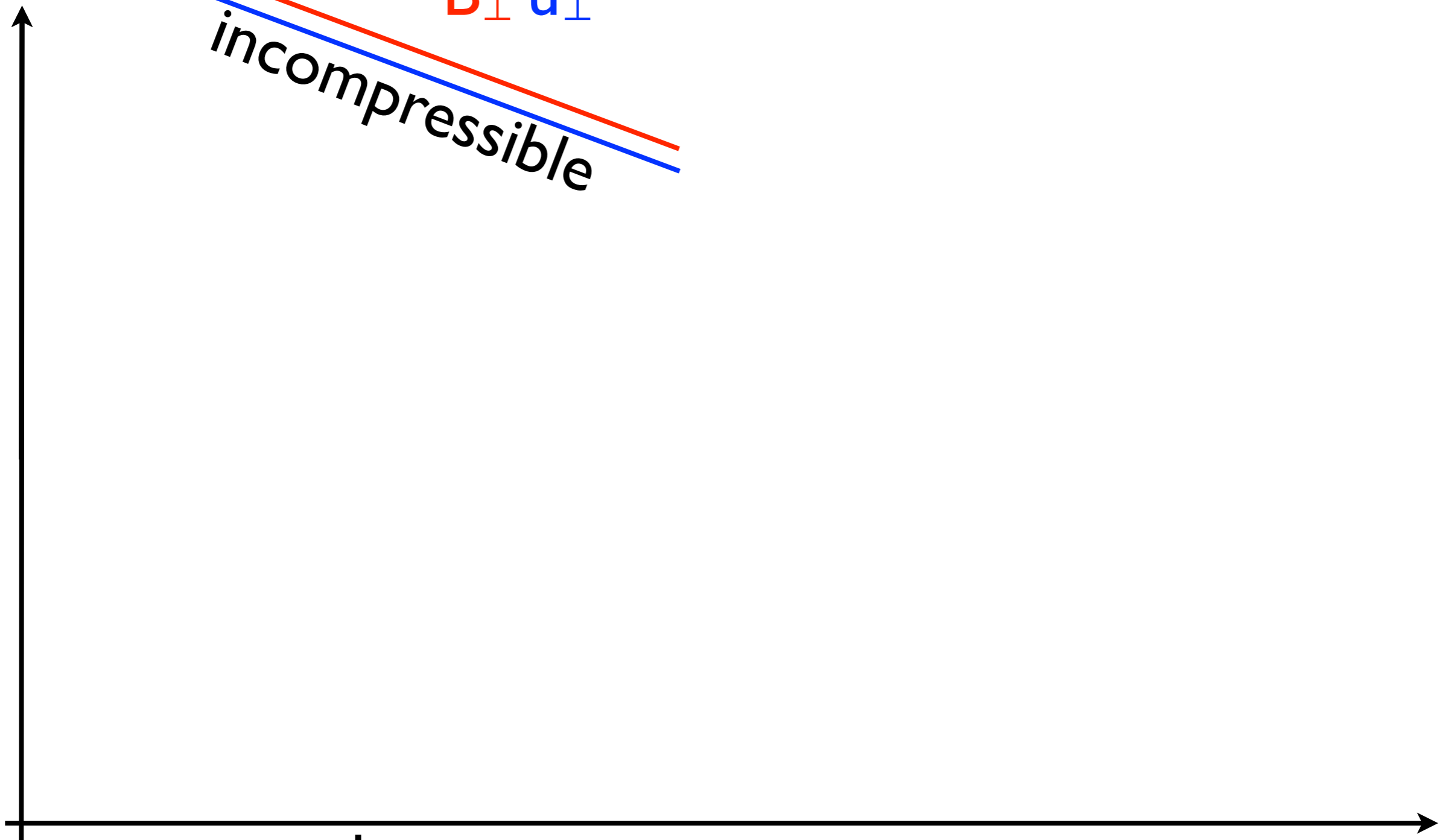


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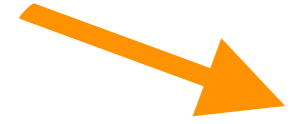
*incompressible*



k

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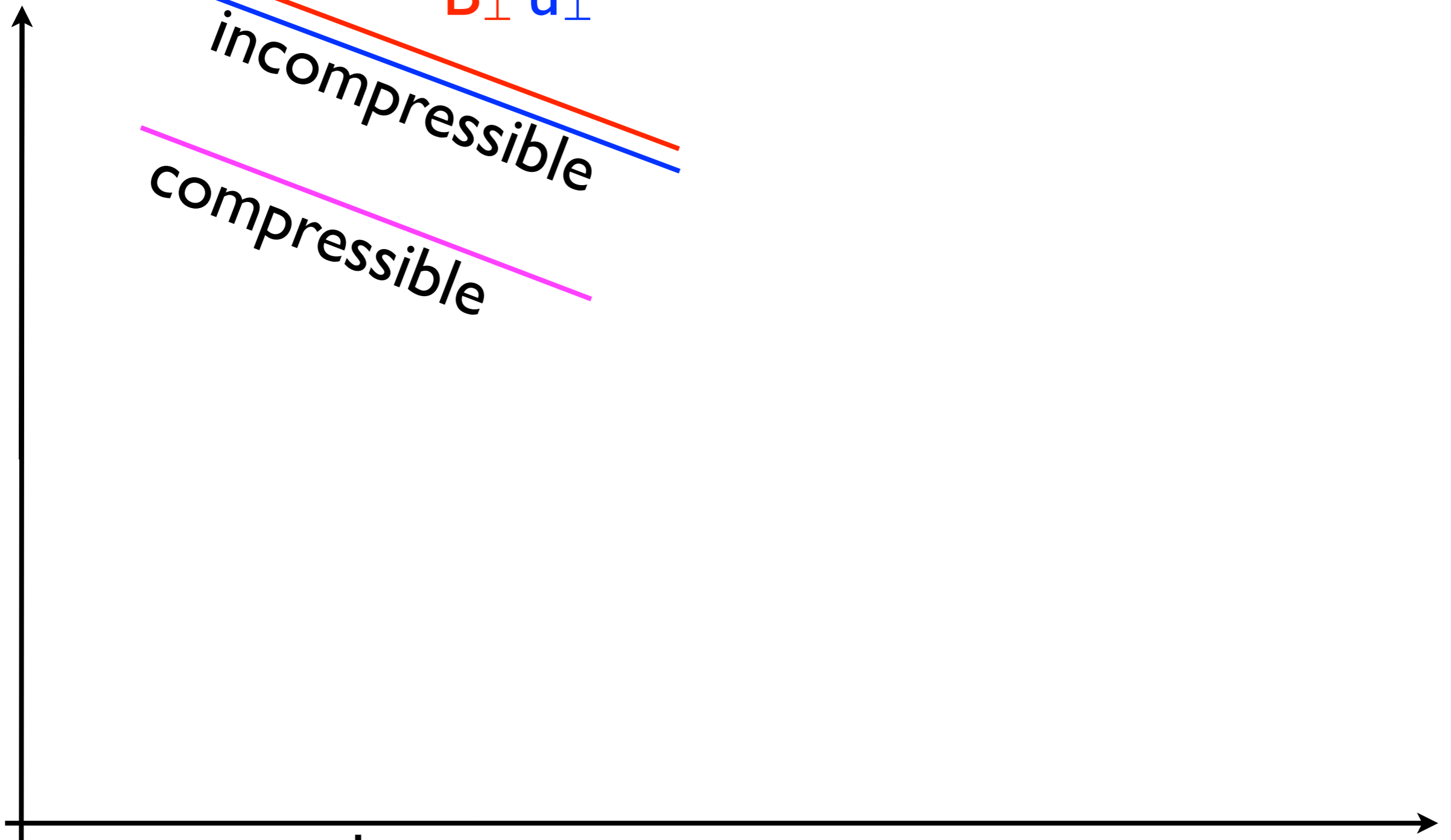


MHD:

$B_{\perp}$   $u_{\perp}$

*incompressible*

*compressible*



k

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Energy



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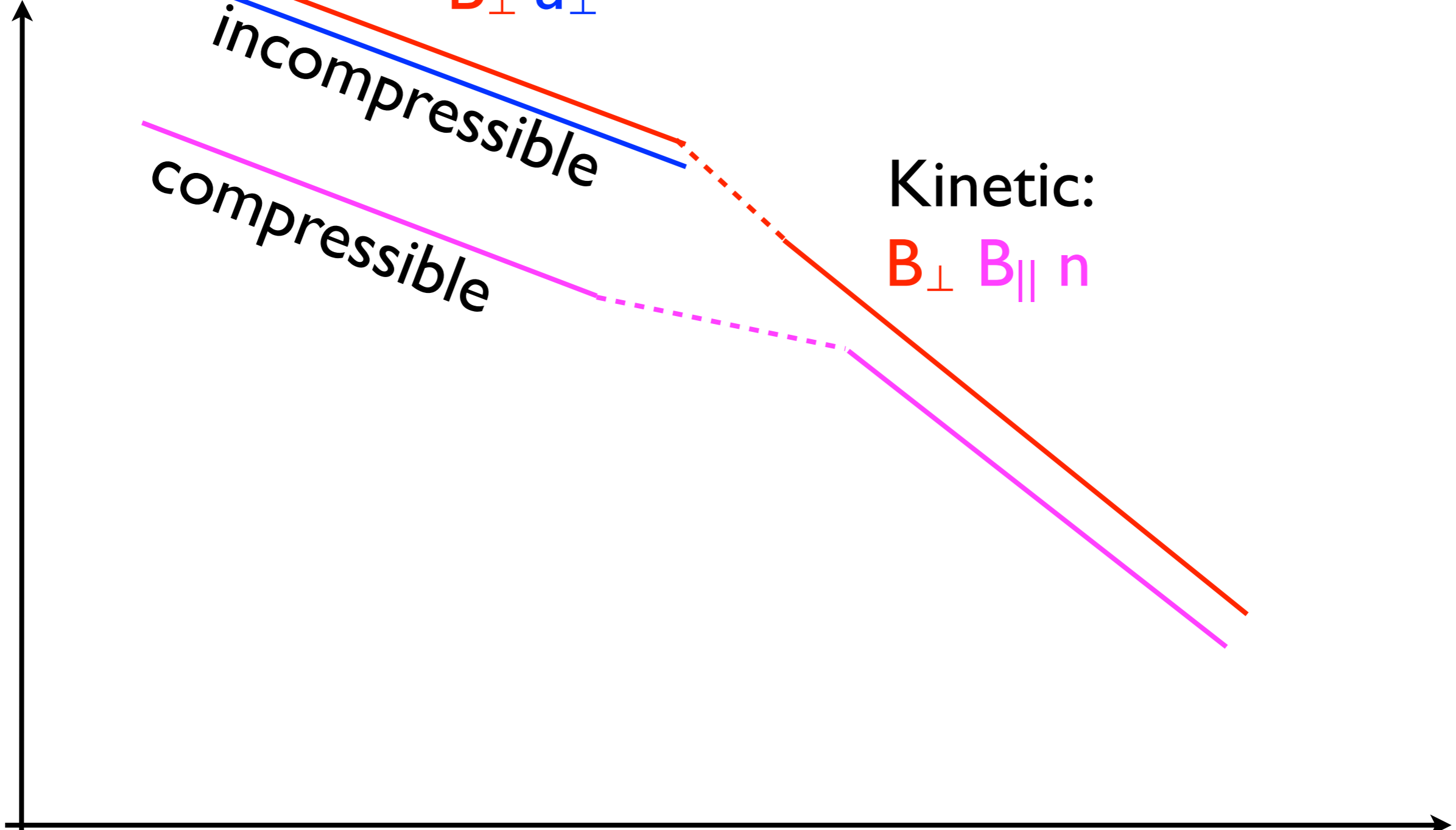
$B_{\perp}$   $u_{\perp}$

*incompressible*

*compressible*

Kinetic:

$B_{\perp}$   $B_{\parallel}$   $n$

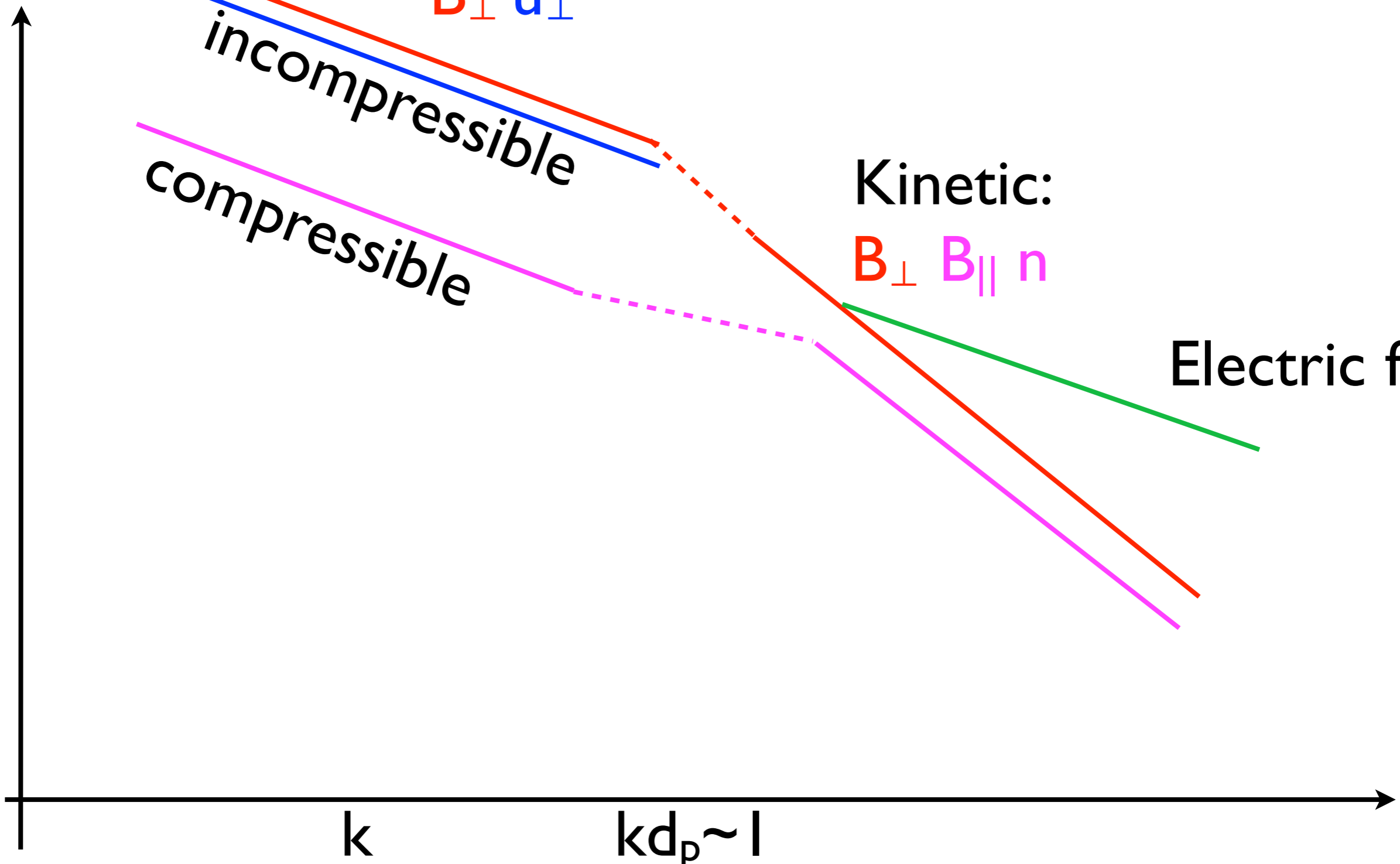


$k$

$kd_p \sim l$

# MHD-kinetic turbulent cascade

Energy



*incompressible*

*compressible*

MHD:  
 $B_{\perp}$   $u_{\perp}$

Kinetic:  
 $B_{\perp}$   $B_{\parallel}$   $n$

Electric field

$k$

$kd_p \sim 1$



# MHD-kinetic turbulent cascade

Energy



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$B_{\perp}$   $u_{\perp}$

*incompressible*

*compressible*

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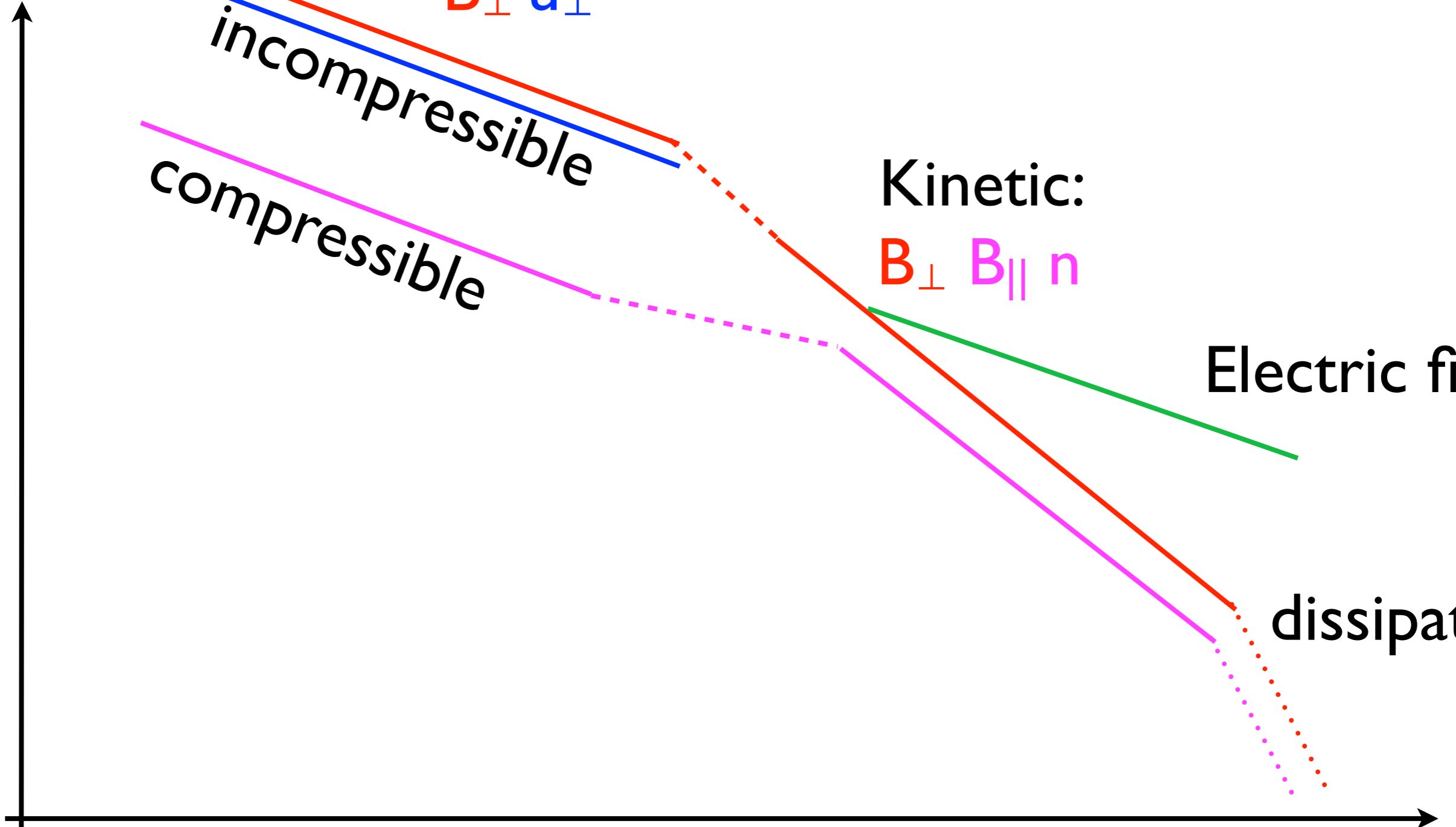
Electric field

dissipation

$k$

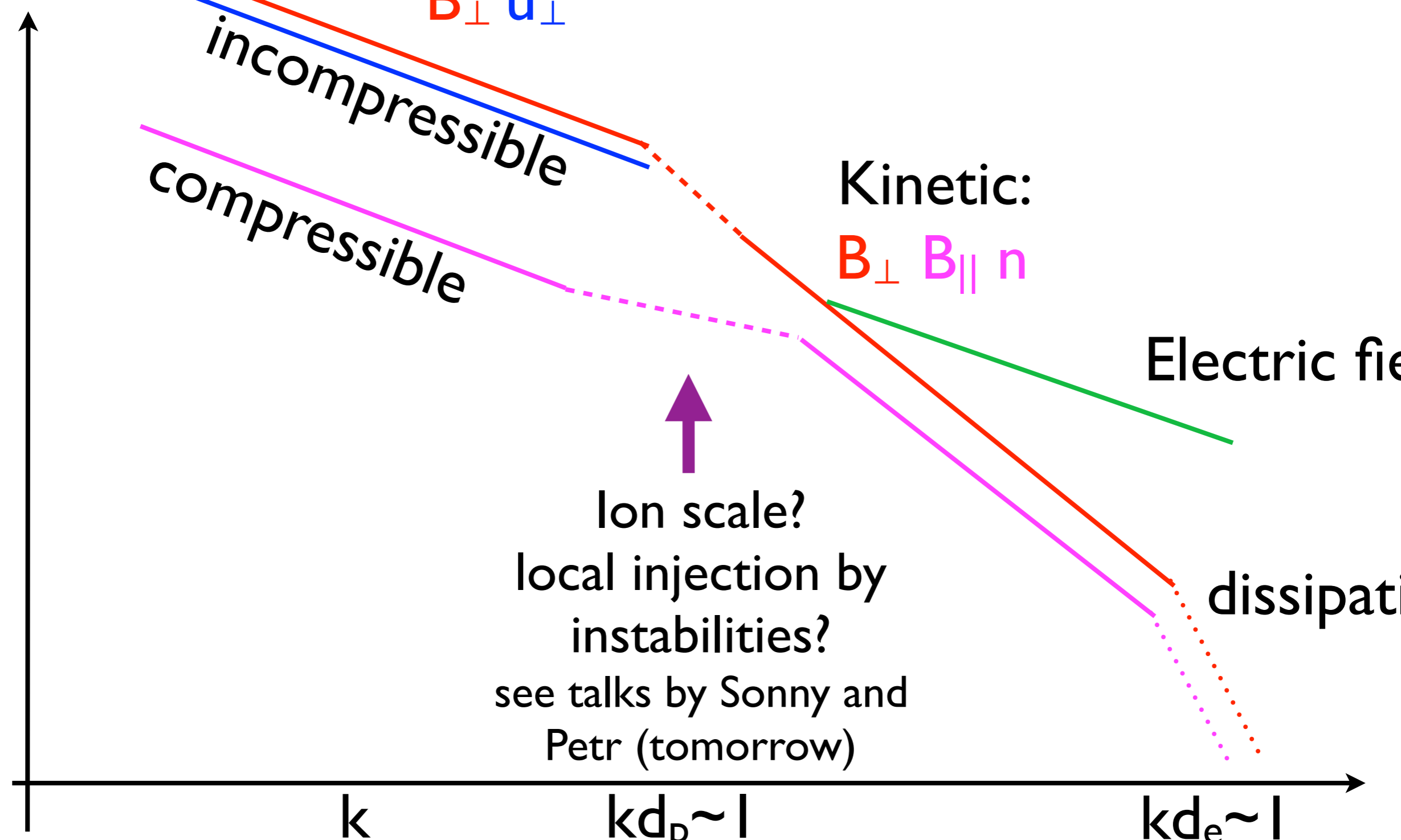
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Energy



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*incompressible*

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Electric field

dissipation



Ion scale?

local injection by instabilities?

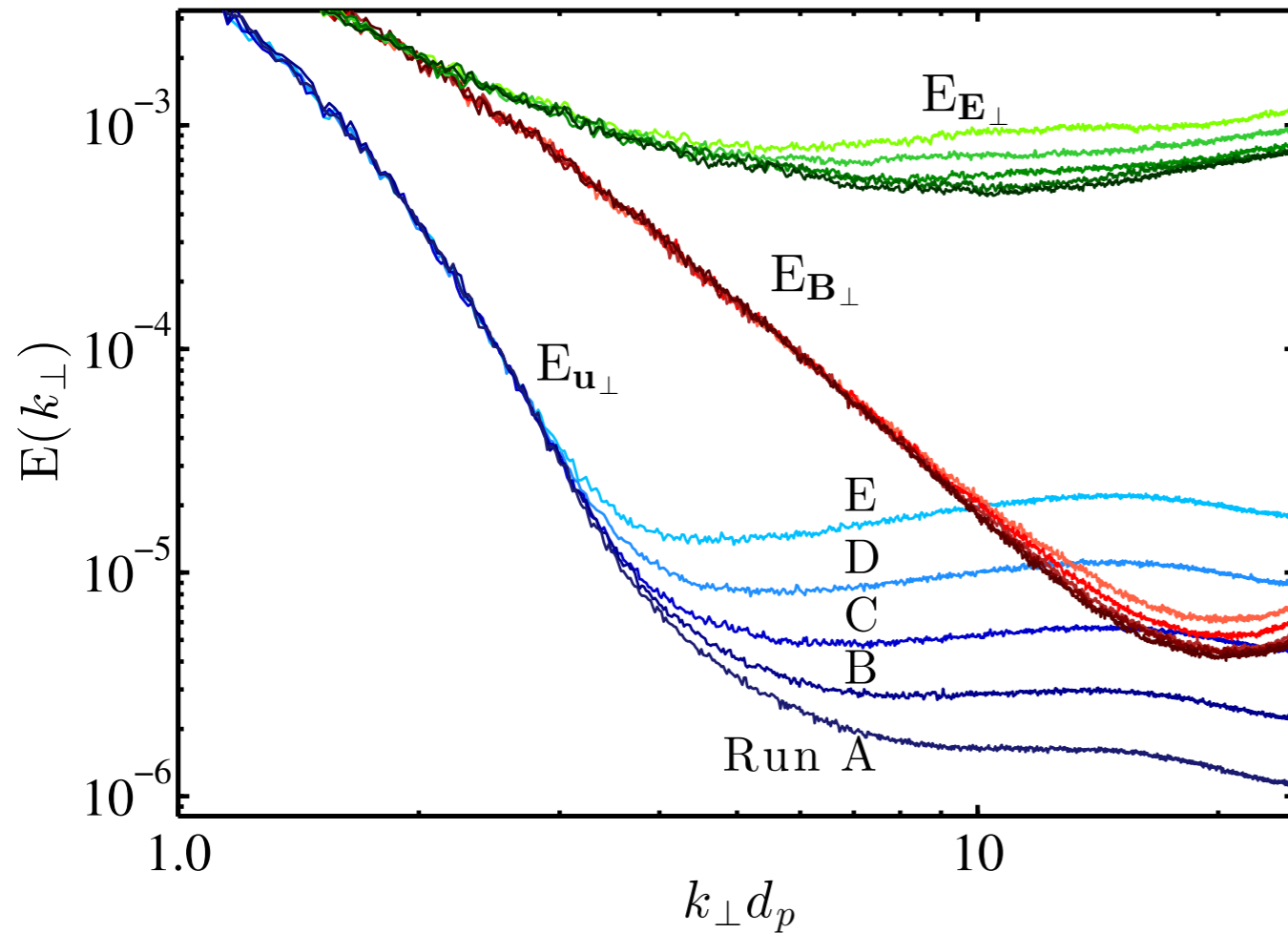
see talks by Sonny and Petr (tomorrow)

$k$

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$kd_e \sim l$

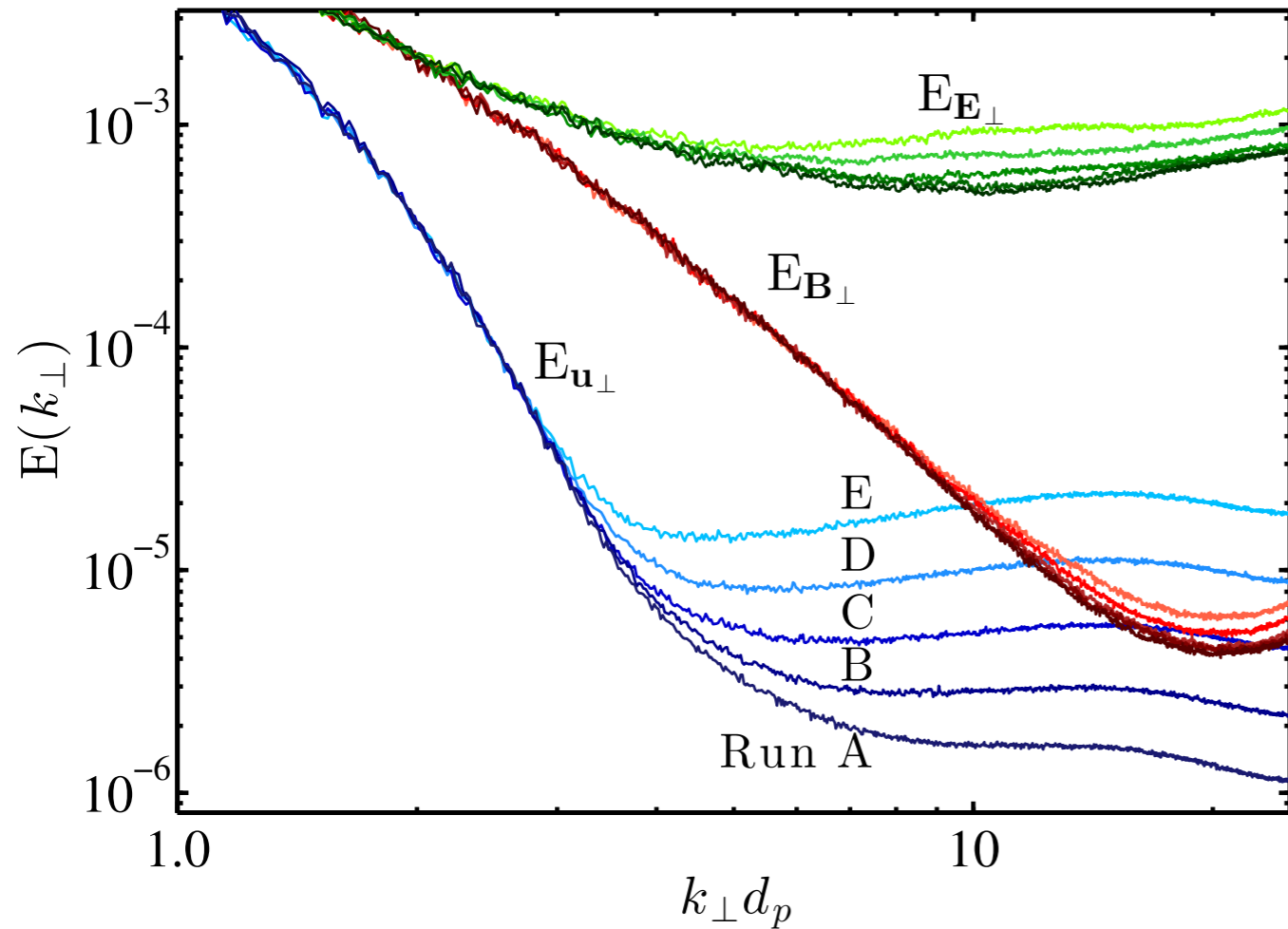
# Stability of results: resistivity and ppc



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Number of particles per cell  
influences the spectral properties  
of different fields

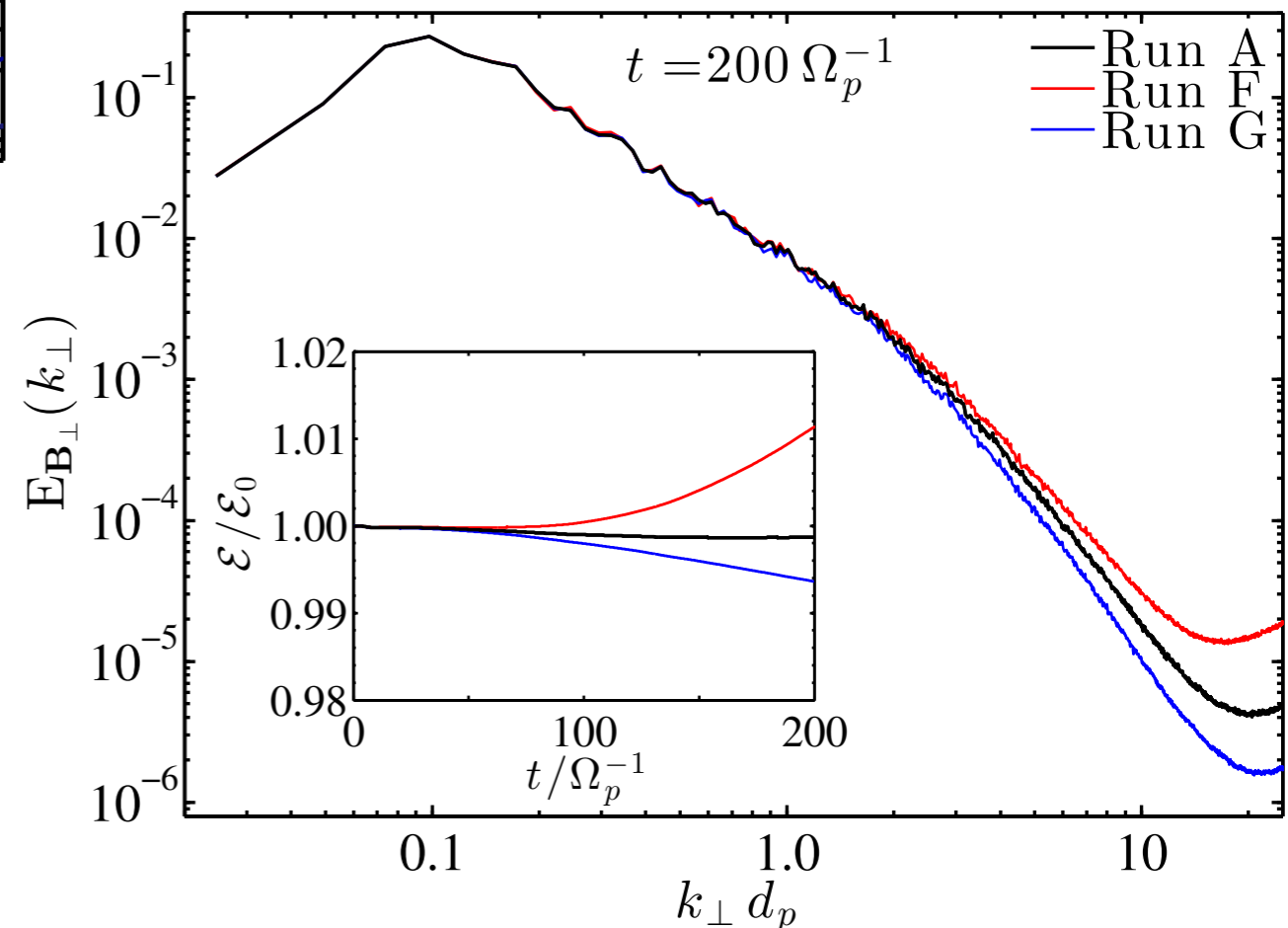
# Stability of results: resistivity and ppc



Number of particles per cell influences the spectral properties of different fields

Run	$\Delta x/d_p$	$L_{\text{box}}/d_p$	$\eta$	ppc
A	0.125	256	$5 \times 10^{-4}$	8000
B	0.125	256	$5 \times 10^{-4}$	4000
C	0.125	256	$5 \times 10^{-4}$	2000
D	0.125	256	$5 \times 10^{-4}$	1000
E	0.125	256	$5 \times 10^{-4}$	500
F	0.125	256	$1 \times 10^{-4}$	8000
G	0.125	256	$1 \times 10^{-3}$	8000
H	0.250	512	$1 \times 10^{-3}$	8000
I	0.500	1024	$2 \times 10^{-3}$	8000

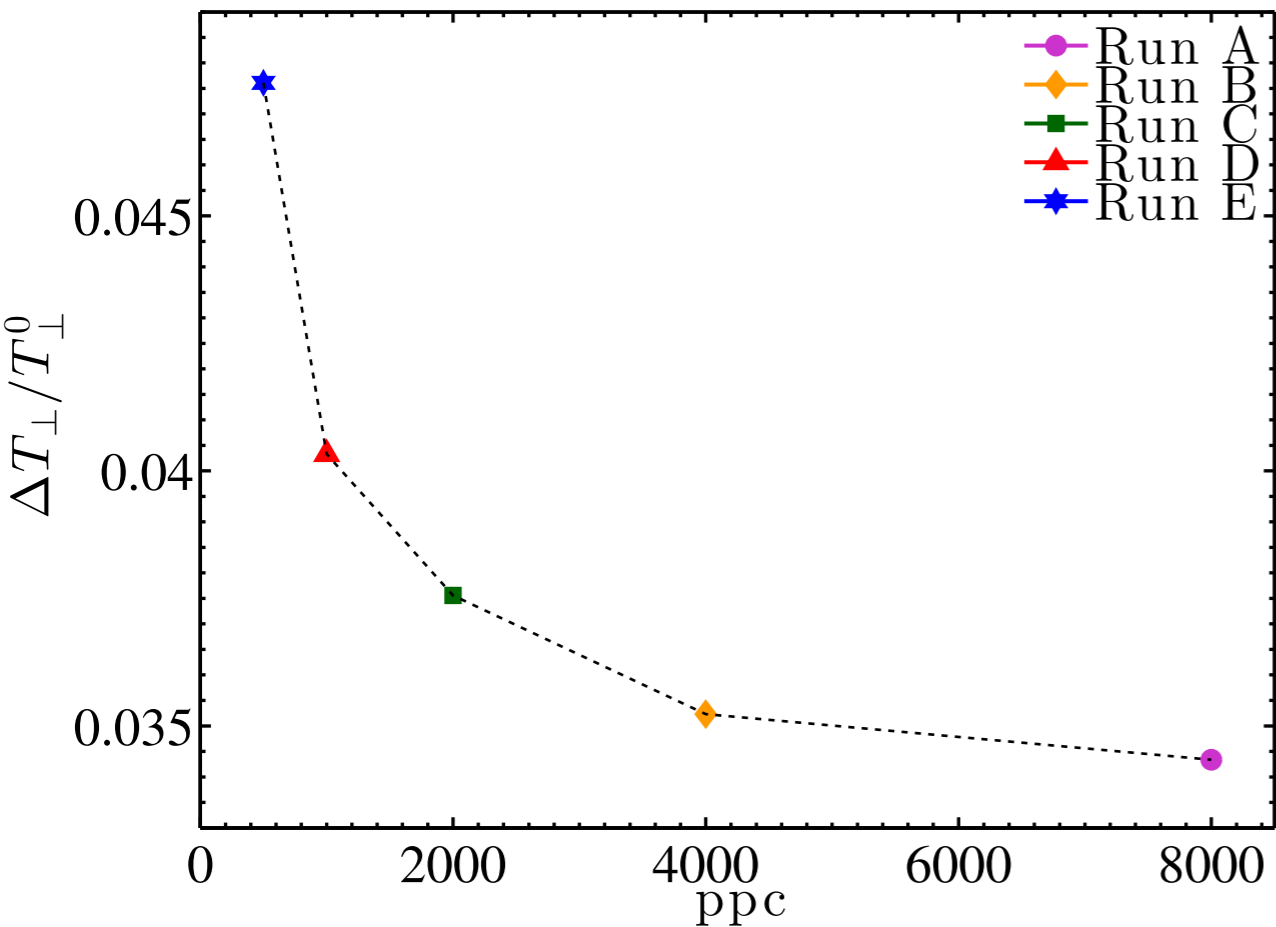
Fine tuning of resistivity is needed to satisfy energy conservation



# Stability of results: particle heating

Number of particles per cell  
influences the amount of  
perpendicular heating.

We see convergence for  $\text{ppc} > 4000$

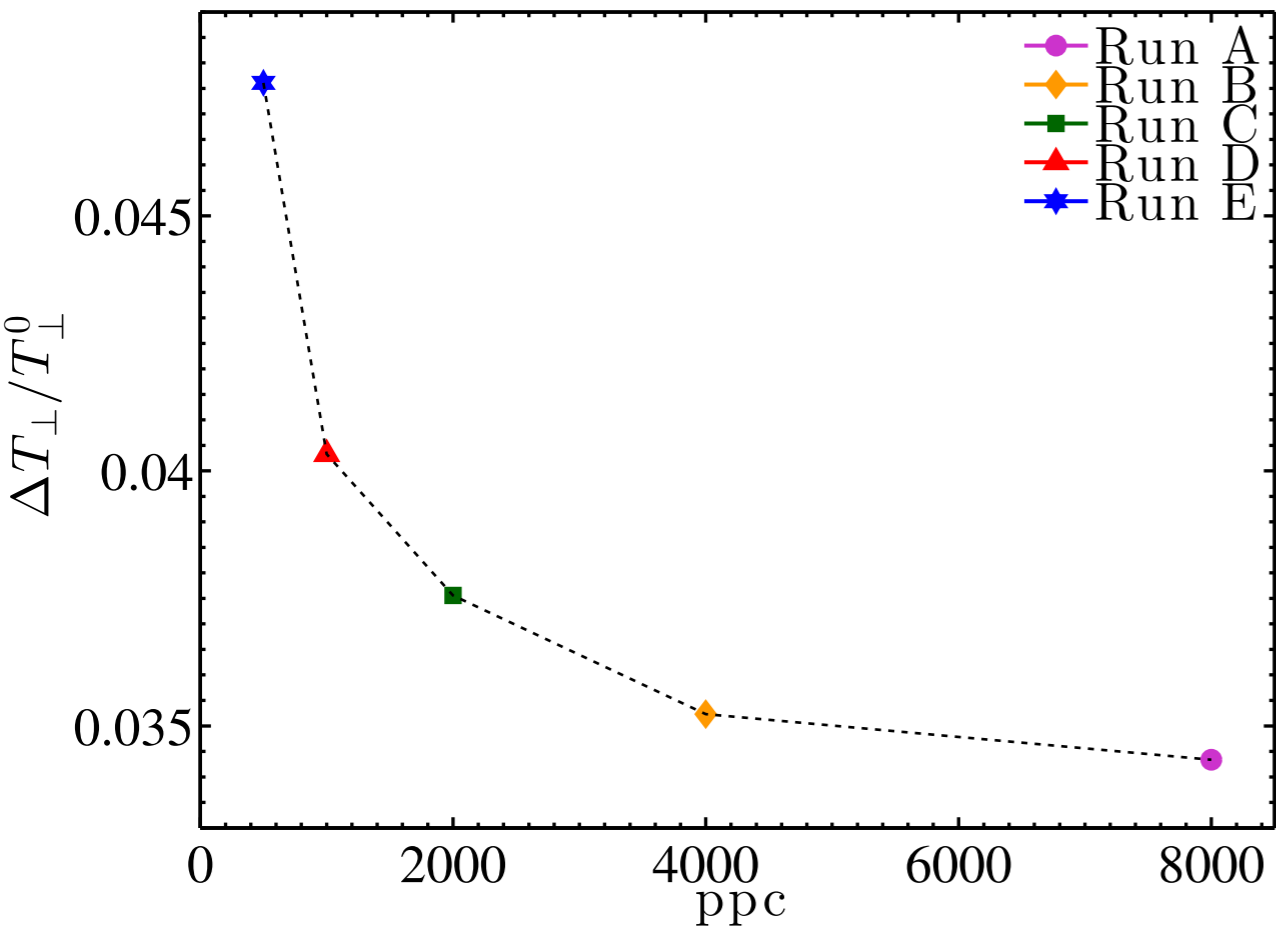


Run	$\Delta x / d_p$	$L_{\text{box}} / d_p$	$\eta$	ppc
A	0.125	256	$5 \times 10^{-4}$	8000
B	0.125	256	$5 \times 10^{-4}$	4000
C	0.125	256	$5 \times 10^{-4}$	2000
D	0.125	256	$5 \times 10^{-4}$	1000
E	0.125	256	$5 \times 10^{-4}$	500
F	0.125	256	$1 \times 10^{-4}$	8000
G	0.125	256	$1 \times 10^{-3}$	8000
H	0.250	512	$1 \times 10^{-3}$	8000
I	0.500	1024	$2 \times 10^{-3}$	8000

# Stability of results: particle heating

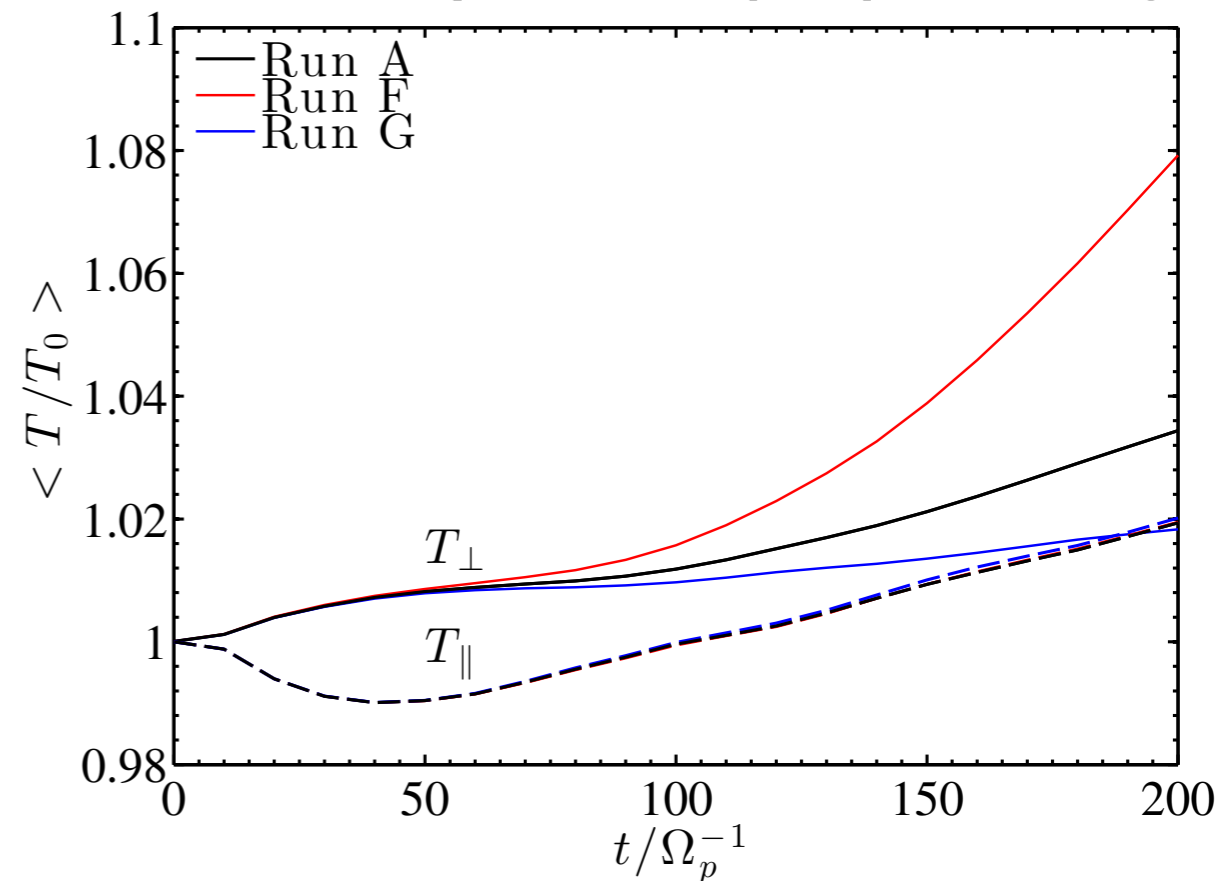
Number of particles per cell  
influences the amount of  
perpendicular heating.

We see convergence for  $\text{ppc} > 4000$



Run	$\Delta x / d_p$	$L_{\text{box}} / d_p$	$\eta$	ppc
A	0.125	256	$5 \times 10^{-4}$	8000
B	0.125	256	$5 \times 10^{-4}$	4000
C	0.125	256	$5 \times 10^{-4}$	2000
D	0.125	256	$5 \times 10^{-4}$	1000
E	0.125	256	$5 \times 10^{-4}$	500
F	0.125	256	$1 \times 10^{-4}$	8000
G	0.125	256	$1 \times 10^{-3}$	8000
H	0.250	512	$1 \times 10^{-3}$	8000
I	0.500	1024	$2 \times 10^{-3}$	8000

Resistivity affects perp. heating

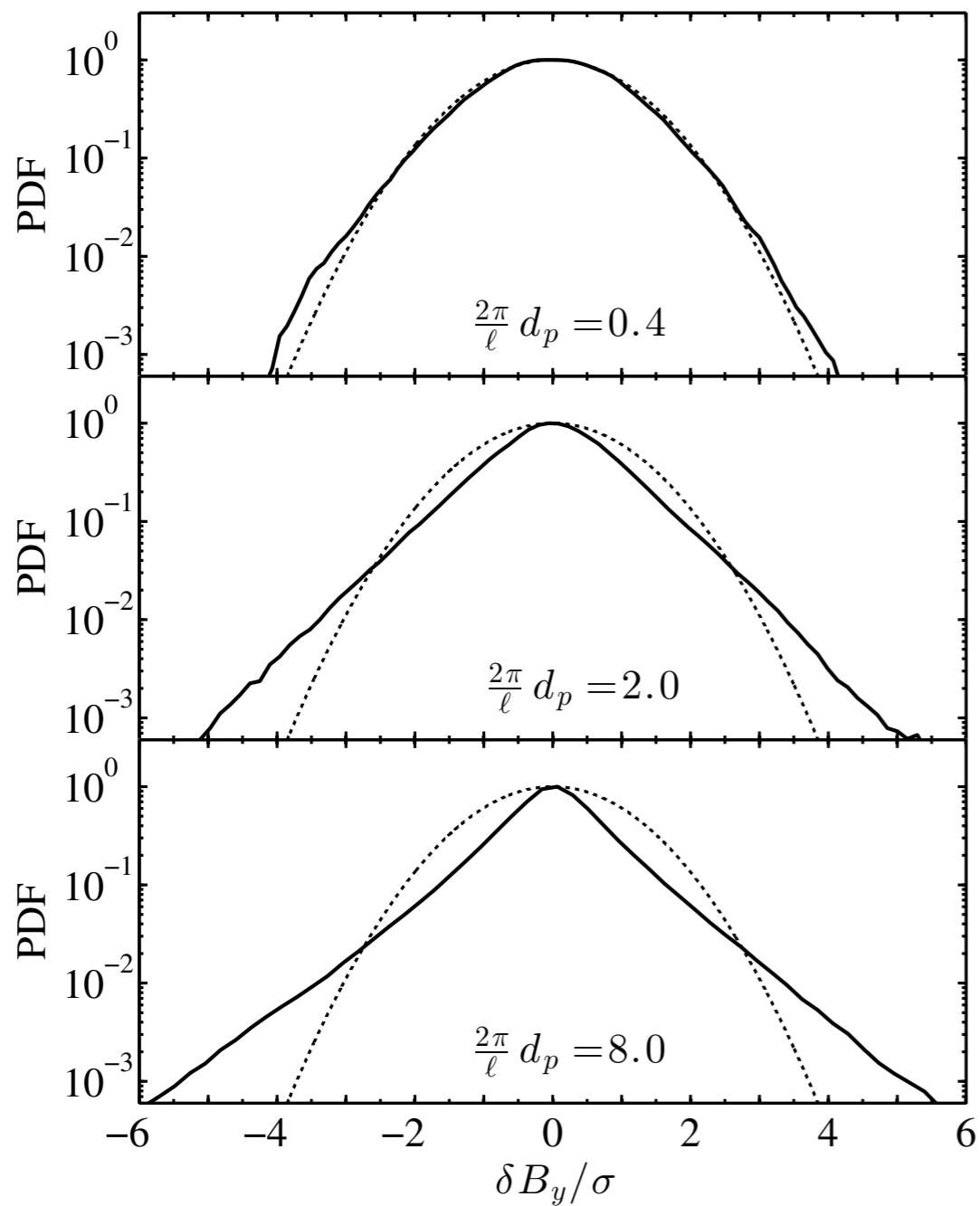


# Conclusion

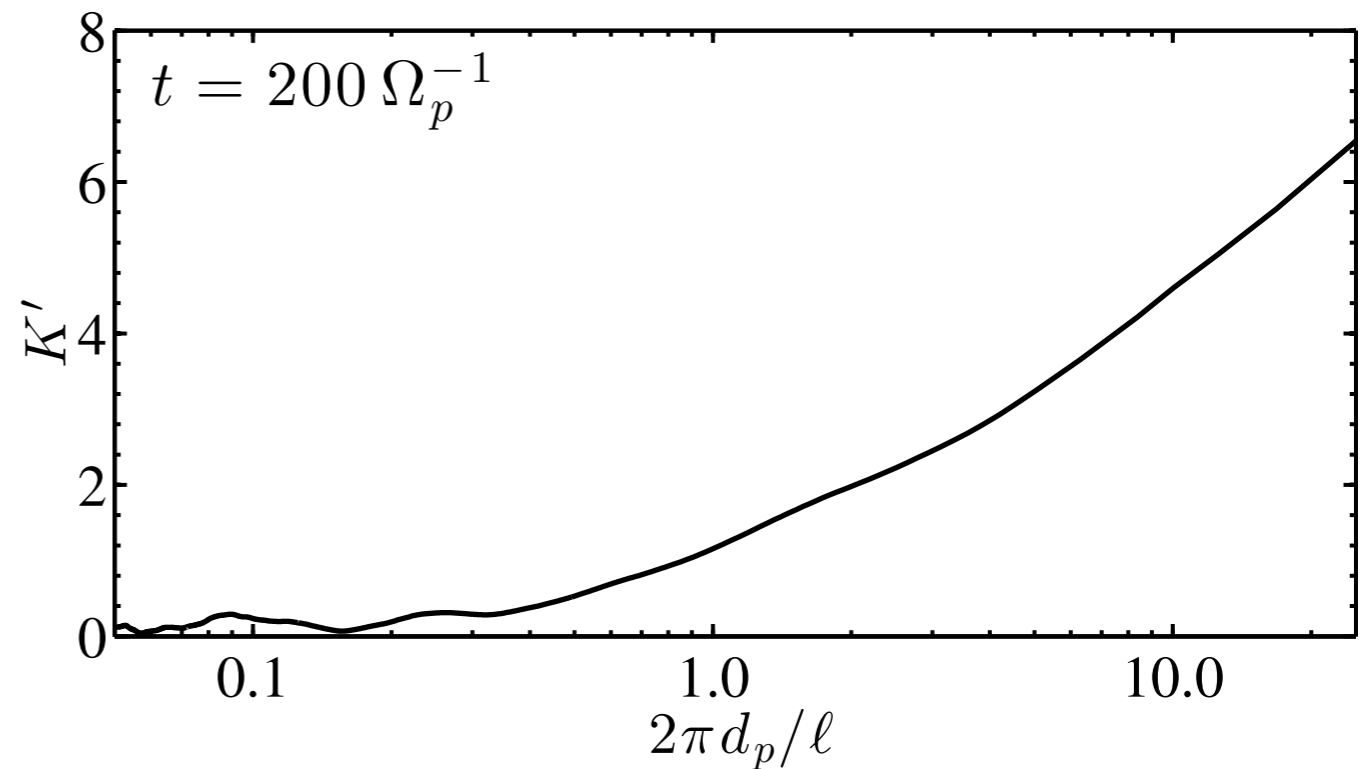
- Self-consistent description of transition from MHD to kinetic turbulence. One full decade for each regime separated by a smooth break.
- 2D geometry can capture most of SW properties. Why??  
3D is needed to validate results.
- MHD with residual energy  $\sim k^{-2}$  (-5/3, -3/2, -2)
- A cascade at kinetic scales consistent with KAW turbulence (fluctuations not waves!)
- Some proton heating in both parallel and perp directions is observed. Parallel one is very solid, perp. heating is strongly affected by resolution. Some care has to be used in general to interpret simulation results!  
More work is needed to understand the origin of the heating

# Intermittency

Non-Gaussian distribution of  $\delta B$  increasing with scale

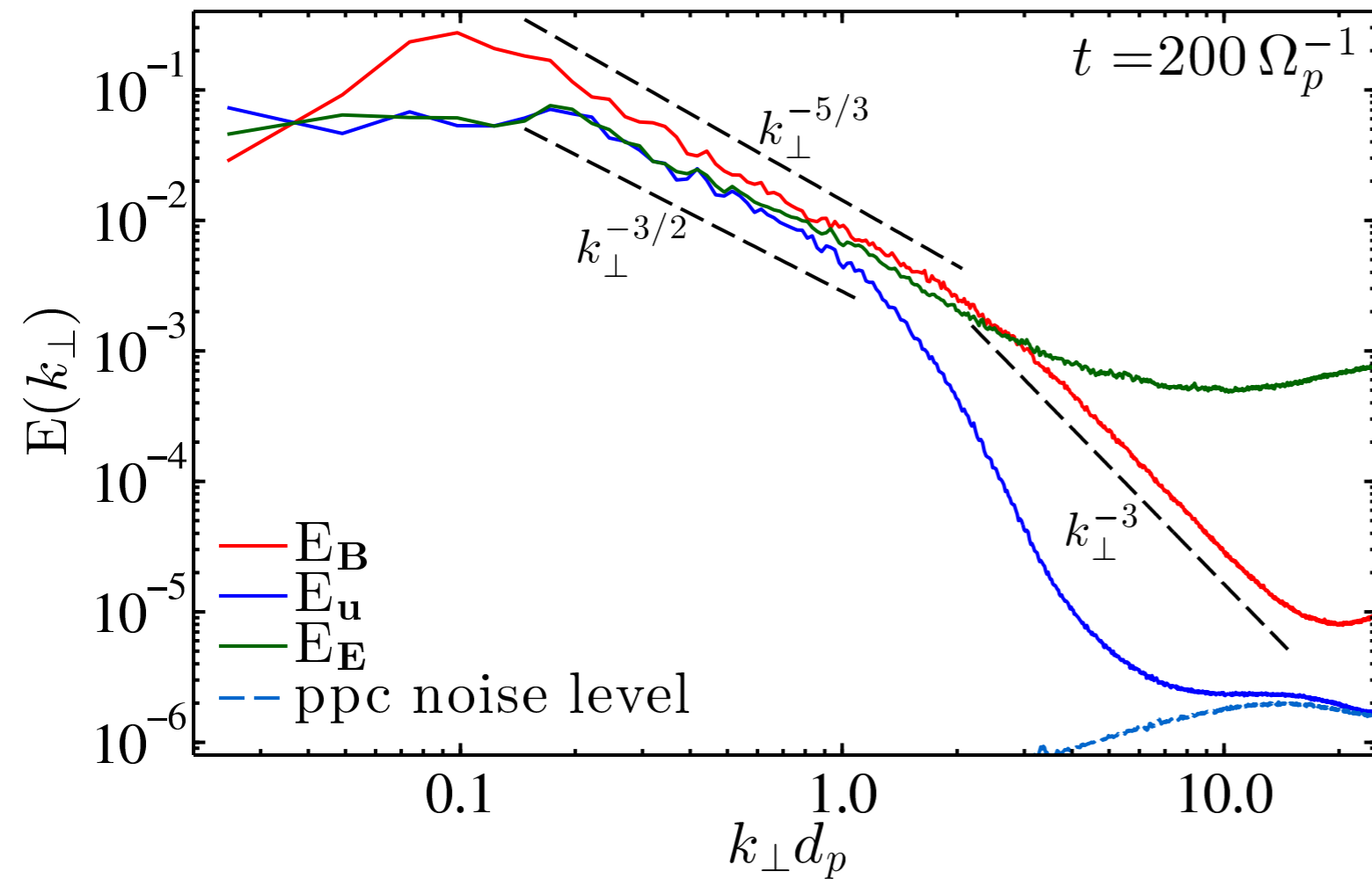


## Excess of Kurtosis





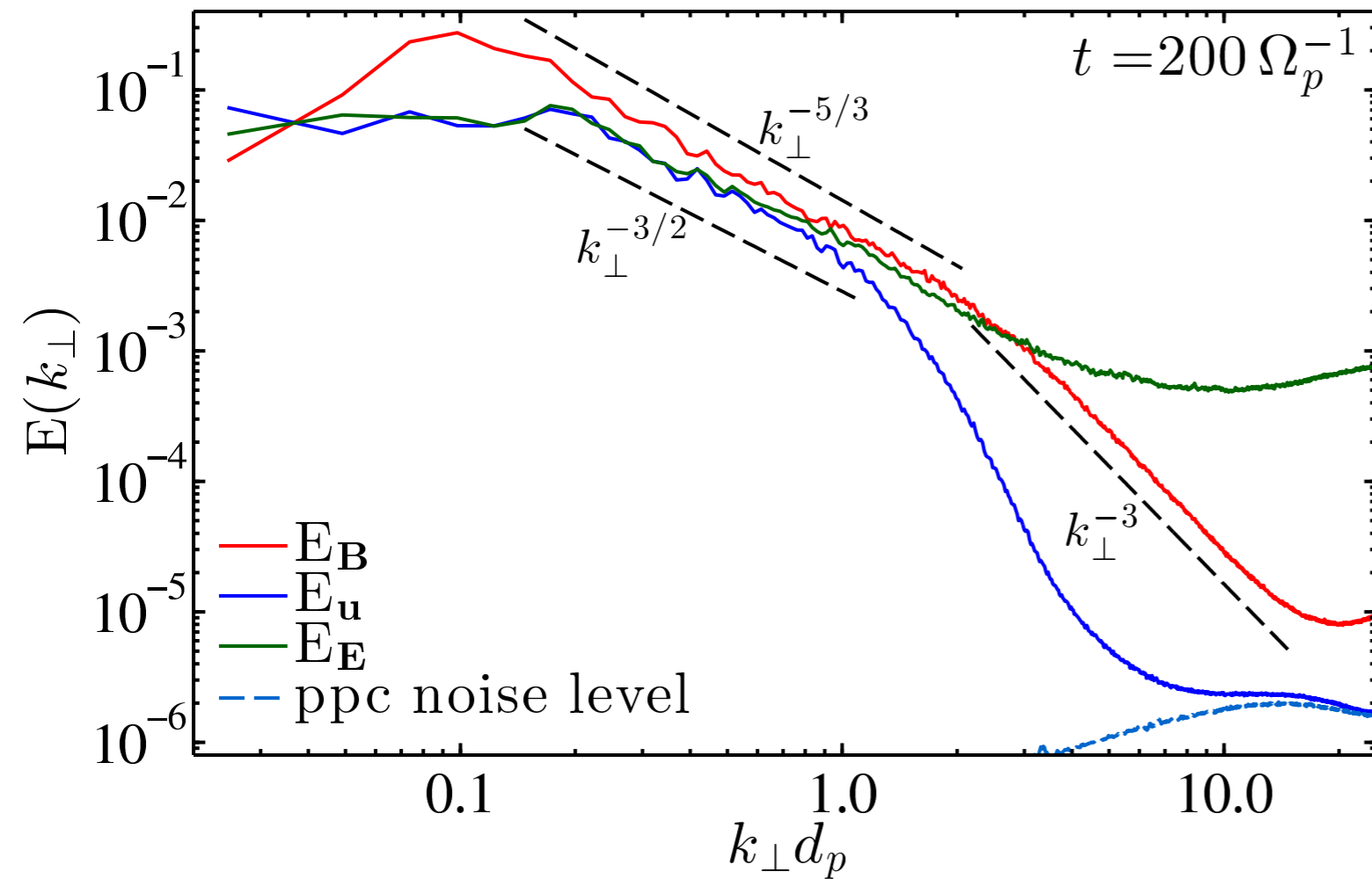
# Electric field (inertial range)



MHD range:  $E = -V \times B \sim \delta v B_0 + \delta b U_0$

Plasma frame:  $E_{SW} = \delta v B_0 \propto k^{-3/2}$

# Electric field (inertial range)



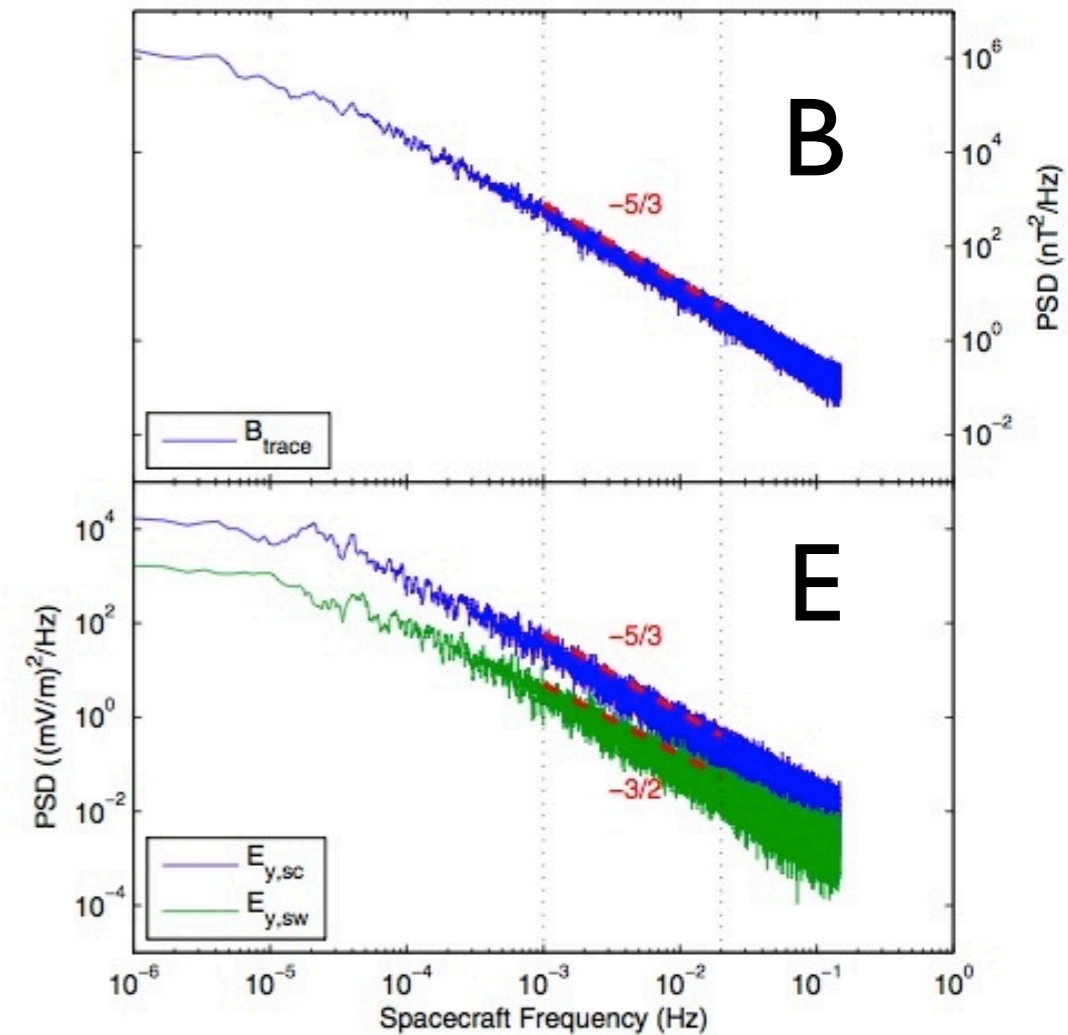
MHD range:  $E = -V \times B \sim \delta v B_0 + \delta b U_0$

Plasma frame:  $E_{sw} = \delta v B_0 \propto k^{-3/2}$

s/c frame:

$E_{s/c} = \delta v B_0 + \delta b U_0 \sim \delta b U_0 \propto k^{-5/3}$

E is frame dependent



**Figure 1.** Sample power spectra from *P2* 2010 days 217–230. The dotted lines show the range of scales that the spectral index is fitted to. Gradients of  $-5/3$  and  $-3/2$  are marked for reference.

(A color version of this figure is available in the online journal.)

Chen et al 2011

# Electric field - all terms

$$\underline{\mathbf{E}} = \underbrace{-\mathbf{u} \times \mathbf{B}}_{\mathbf{E}_{\text{MHD}}} + \underbrace{\mathbf{J} \times \mathbf{B}/n}_{\mathbf{E}_{\text{Hall}}} - \underbrace{\nabla p_e/n}_{\mathbf{E}_{\text{pe}}} + \underbrace{\eta \mathbf{J}}_{\mathbf{E}_\eta}$$

